- A fast track towards the spacetime geometry? -

C03 : Masahiro Ibe [ ICRR & IPMU ] 8/31/2013 @ Neutrino Frontier Workshop

Based on Phys.Rev.D86(2012)013002 (K.Harigaya, MI, T.T.Yanagida)

✓ Why are we so much interested in neutrino mass?

Neutrino mass is a window to high energy physics beyond the Standard Model !

- ✓ Tiny ! → New mass scales ? New symmetry ?
- ✓ Mixing ! → Implications on flavor structure ?
- Majorana ?  $\rightarrow$  Lepton number violation ?
- CP-violation ?  $\rightarrow$  Baryon asymmetry of the universe ?

Seesaw Mechanism ['79 Yanagida; '79 Gell-Mann, Ramond, Slansky]

In the Standard Model :

$$\mathcal{L} = y_{\alpha\beta} \ell_{L\alpha} \bar{e}_{R\beta} h \qquad \langle h \rangle = v \simeq 174.1 \,\text{GeV}$$
  
(\alpha, \beta = \eta, \mu, \tau) \qquad \text{the neutrinos remain massless !

Let us introduce the right-handed neutrinos (N<sub>i</sub>):

$$\mathcal{L} = y_{\alpha\beta}\ell_{L\alpha}\bar{e}_{R\beta}h + \lambda_{i\alpha}N_{i}\ell_{L\alpha}h - \frac{1}{2}M_{ij}N_{i}N_{j}$$

$$\rightarrow \mathcal{L}_{mass}^{\nu} = -\frac{1}{2}\left[\left(\nu_{L}, N_{R}\right)\begin{pmatrix}0 & m_{D}^{T}\\m_{D} & M\end{pmatrix}\left(\begin{array}{c}\nu_{L}\\N_{R}\end{array}\right)\right] + h.c.$$

$$\rightarrow \text{the neutrinos have finite masses}: \quad m_{\nu} \simeq \frac{m_{D}m_{D}^{T}}{M}$$

 $m_v = O(0.01) \text{ eV for } M = O(10^{11}) \text{ GeV } \& m_D = O(1) \text{GeV } !$ 

### Introduction

Leptogenesis [`86 Fukugita & Yanagida]

**Baryon asymmetry** (from nucleosynthesis and CMB):

$$\eta_{B_0} = \frac{n_B - n_{\bar{B}}}{n_\gamma} \simeq 6 \times 10^{-10}$$

must have been generated during the evolution of the universe.

Sakharov three conditions ('67) :

- ✓ B (or B-L) symmetry breaking
- C and CP violation

B-L and C/CP violating interactions outside of thermal equilibrium

## Introduction

Leptogenesis [`86 Fukugita & Yanagida]

Time

Inflation :  $T \rightarrow 0$ ,  $\eta_{B0} \rightarrow 0$ Sakharov conditions Reheating:  $T \rightarrow T_R$ ,  $\eta_{B0} = 0$ 1: N<sub>R</sub> mass violates L  $N_R \rightarrow \ell + h, \, \ell^* + h^*$  $N_R$  are in the thermal equilibrium ( $T \gg M_R$ ) 2: CP-violating decay  $N_R$  decays at a temperature  $T_D$  $\Gamma[N_R \to \ell + h]$  $\neq \Gamma[N_R \rightarrow \ell^* + h^*]$ Lepton asymmetry is generated ! 3: Out of equillibrium  $M_R/T_D \gtrsim 1$  $\langle n_L \rangle \neq 0 \longrightarrow Sphaleron \longrightarrow \langle n_B \rangle \neq 0$ Baryon asymmetry is generated !  $\eta_{B0} \simeq 3 \times 10^{-10} \times \left(\frac{M_R}{10^{10} \,\mathrm{GeV}}\right) \left(\frac{m_{\nu}^{\mathrm{eff}}}{0.05 \mathrm{eV}}\right) \bar{\kappa} \sin \delta_{\mathrm{eff}}$  $\bar{\kappa} \simeq \left(\frac{0.01 \,\mathrm{eV}}{\tilde{m}_1}\right)^{1.16} \left[\tilde{m}_1 = \sum |\lambda_{1\alpha}|^2 \frac{v^2}{M_B} \propto \frac{T_D^2}{M_P^2}\right]$ 

### In the seesaw mechanism...

Tiny neutrino mass can be explained by a new scale
= Right handed neutrino mass !

With the CP-violating phases in the right-handed neutrino sector, the Baryon Asymmetry of the universe can be explained by Leptogenesis.

Future observations of the *CP*-asymmetry in the neutrino oscillations and the neutrino-less double beta decay, will support the ideas of the *seesaw mechanism* and *Leptogenesis* qualitatively.

To what extent will we learn the seesaw mechanism and Leptogenesis quantitatively?

# Seesaw Mechanism vs Neutrino oscillation



For given  $\overline{m}_{vi}$  and  $U_{MNS}$  in the seesaw mechanism

$$\bar{m}_{\nu} = U_{MNS}^T \lambda^T M_R^{-1} \lambda U_{MNS} v^2$$

the Yukawa coupling  $\lambda$  is determined up to R,

$$\lambda = \frac{1}{v} M_R^{1/2} R \, \bar{m}_{\nu}^{1/2} \, U_{MNS}^{\dagger}$$

which satisfies  $R^{T}R = 1$  (i.e. complex orthogonal matrix = 6 parameters).

The Yukawa coupling  $\lambda$  cannot be determined by the low energy data...

Relation between CP-violating phases :

Neutrino oscillation : Dirac CP-phase  $\delta$  in  $U_{MNS}$ 

 $A_{CP} = P(v_{l} \rightarrow v_{l'}) - P(\bar{v}_{l} \rightarrow \bar{v}_{l'}) \propto J_{CP} = Im[U_{\mu3}U_{e3}^{*}U_{e2} U_{\mu2}^{*}]$  $= (sin2\theta_{12} sin2\theta_{12} sin2\theta_{13} cos\theta_{13} sin\delta)/8$ 

### Leptogenesis : CP-phase of the redundant parameters in R

$$\eta_{B0} \propto m_{\nu}^{\text{eff}} \sin \delta_{\text{eff}} \qquad m_{\nu}^{\text{eff}} \sin \delta_{\text{eff}} = \frac{\text{Im}[\lambda m_{\nu} \lambda^{T}]_{11}}{(\lambda \lambda^{\dagger})_{11}}$$
$$\lambda \lambda^{\dagger} = \frac{1}{v^{2}} M_{R}^{1/2} R \bar{m}_{\nu} R^{\dagger} M_{R}^{1/2} \qquad \lambda m_{\nu} \lambda^{T} = \frac{1}{v^{2}} M_{R}^{1/2} R \bar{m}_{\nu}^{2} R^{T} M_{R}^{1/2}$$
$$\rightarrow \eta_{B0} \text{ does not depend on } U_{MNS} \dots$$

The *CP*-violating phases in the neutrino oscillation and Leptogenesis are independent.

# Seesaw Mechanism vs Neutrino oscillation

- The seesaw mechanism is attractive model to explain the observed tiny neutrino mass.
- Without knowing the origin of λ, it is difficult to test the seesaw mechanism from the low energy data.
- Observation of the CP-asymmetry in neutrino oscillations will support Leptogenesis qualitatively, but they are quantitatively independent.

### To go one step further?

Top down : Flavor symmetries, Grand Unified Theory...

Instead, we take a **bottom up** approach as a trial where we reduce the number of the Yukawa couplings as small as possible as long as the experimental results are reproduced (Occam's Razor).

#### We need only two right-handed neutrinos!

 $\bar{m}_{\nu} = U_{MNS}^{T} \lambda^{T} M_{R}^{-1} \lambda U_{MNS} v^{2}$ (rank[ $\bar{m}_{\nu}$ ] = min[rank[ $U_{MNS}$ ], rank[ $\lambda$ ], rank[ $M_{R}$ ]])

 $\rightarrow$  the lightest neutrino mass = 0!

and a range the area

Number of real valued parameters

Seesaw Mechanism $M_i$ 2 $y_{\alpha\beta}$ 3 $\lambda_{i\alpha}$ 9 = (12-3)

Low energy theory		
Mi	2	
<b>У</b> αβ	3	
$\overline{m}_{vi}$	2	
U <sub>MNS</sub>	5 = 3 + 1 + 1	

A complex redundant parameter z :

[Normal Hierarchy :  $\overline{m}_{v1} = 0$ ]

$$R = \left(\begin{array}{ccc} 0 & \cos z & -\sin z \\ 0 & \sin z & \cos z \end{array}\right)$$

[Inverted Hierarchy:  $\overline{m}_{v3} = 0$ ]  $R = \begin{pmatrix} -\sin z & \cos z & 0 \\ \cos z & \sin z & 0 \end{pmatrix}$ 

Minimal Yukawa Structure ? (in diagonalized mass bases)



 $\rightarrow$  we have non-trivial predictions on  $U_{MNS}$  and  $\overline{m}_{vi}$ .

['02 Frampton, Glashow, Yangagida, '02 Raidal, Strumia, '04 Ibarra, Ross ]

Do they reproduce the observed 5 parameters?

Mass differences :

$$\Delta m_{21}^2 = 7.59^{+0.20}_{-0.18} \times 10^{-5} \,\mathrm{eV}^2 \,, \qquad \Delta m_{31}^2 = 2.45^{+0.09}_{-0.09} \times 10^{-3} \,\mathrm{eV}^2 \,(NH) \,,$$
$$\Delta m_{31}^2 = -2.34^{+0120}_{-0.09} \times 10^{-3} \,\mathrm{eV}^2 \,(IH) \,,$$

Mixing Angle :

$$\sin^2 \theta_{12} = 0.312^{+0.017}_{-0.015} , \qquad \sin^2 \theta_{23} = 0.51^{+0.06}_{-0.06} (NH) , \quad \sin^2 \theta_{13} = 0.023^{+0.004}_{-0.004} , \\ \sin^2 \theta_{23} = 0.52^{+0.06}_{-0.06} (IH) ,$$

['11 Schwetz, M. Tortola and J. W. F. Valle, '12 Daya Bay]

#### We put two-zeros in $\lambda$

Redundant parameter "z" is fixed. Two relations on  $U_{MNS}$  and  $\overline{m}_{vi}$ .

 $\rightarrow$  5 (out of 7) parameters remain in  $U_{MNS}$  and  $\overline{m}_{vi}$ !

We have sufficient parameters!





A bit small  $sin\theta_{13}$  is predicted...  $\rightarrow$  excluded !

Similarly, all the other possibilities in the normal hierarchy are not consistent with the observed 5 parameters...

For the normal hierarchy with  $m_1 = 0$ , the Yukawa coupling  $\lambda$  depends on  $U_{\alpha 3}$ , and two-zero conditions lead to a sharp prediction on  $sin\theta_{13}$ , which contradicts with observations.

Explicit Yukawa coupling in the normal hierarchy

$$\lambda_{1\alpha} = \frac{1}{v} \sqrt{M_1} \left( \sqrt{m_2} U_{\alpha 2}^* c_z - \sqrt{m_3} U_{\alpha 3}^* s_z \right) ,$$
  
$$\lambda_{2\alpha} = \frac{1}{v} \sqrt{M_2} \left( \sqrt{m_2} U_{\alpha 2}^* s_z + \sqrt{m_3} U_{\alpha 3}^* c_z \right) ,$$





This relation is consistent with data only for  $\delta \simeq \pm \pi/2$  !

In the inverted hierarchy, we found four consistent possibilities :

$$\lambda_{e2} = \lambda_{\mu 1} = 0 \ (\lambda_{e1} = \lambda_{\mu 2} = 0) \qquad \lambda_{e2} = \lambda_{\tau 1} = 0 \ (\lambda_{e1} = \lambda_{\tau 2} = 0)$$

In these cases, we have very sharp predictions !



The effective Majorana neutrino mass

$$m_{ee} = |m_1 U_{e1}^2 + m_2 U_{e2}^2 + m_3 U_{e3}^2$$

In passing...



 $\begin{cases} \delta \simeq \pi/2 & \text{is getting excluded...} \\ \delta \simeq -\pi/2 & \text{is getting favored...??} \end{cases}$ 

Implications on Leptogenesis

**Neutrino oscillation** : Dirac *CP*-phase  $\delta$  in  $U_{MNS}$ **Leptogenesis** : *CP*-phase of the *z* in *R* 

They are now interrelated !

$$\eta_{B0} \propto m_{\nu}^{\text{eff}} \sin \delta_{\text{eff}} = \frac{\Delta m_{12}^2}{\tilde{m}_1} \text{Im}[c_z^2] \qquad \tilde{m}_1 = (\lambda \lambda^{\dagger})_{11} \frac{v^2}{M_R}$$
$$\text{Im}[c_z^2] = \pm s_{12} c_{12} t_{23} s_{13} \sin \delta = \pm \frac{J_{CP}}{c_{13}^2 c_{23}^2}$$
$$(\text{plus}): \lambda_{e_1} = \lambda_{\mu_2} = 0, \lambda_{e_1} = \lambda_{\tau_2} = 0 \qquad (\text{minus}): \lambda_{e_2} = \lambda_{\mu_1} = 0, \lambda_{e_2} = \lambda_{\tau_1} = 0$$

The observation of the *CP*-violation in the neutrino oscillation directly probe the *CP*-violation in Leptogenesis!

$$\eta_{B_0} \simeq \pm 5.9 \times 10^{-10} \times \left(\frac{M_1}{5 \times 10^{13} \,\mathrm{GeV}}\right)$$

- The seesaw mechanism is an attractive framework which explains the tiny neutrino masses!
- The seesaw mechanism also makes it possible to explain the Baryon Asymmetry of the universe via Leptogenesis.
- The seesaw mechanism does not give any particular predictions on the mixing angles and the masses...
- The CP-violation used in Leptogenesis is independent from the CP-violation in the neutrino oscillations...

In the spirit of the Occam's Razor, it is possible to reduce the seesaw mechanism down to...

> Two right-handed neutrino Two zeros in the Yukawa coupling  $\lambda$ .

### Summary

Once the seesaw mechanism is *shaved* down to this level...

Surprisingly sharp predictions !

- One massless neutrino

Inverted hierarchy!  $\delta \simeq \pm \frac{\pi}{2}$   $m_{ee} \simeq 47 \,\mathrm{meV}$ 

The CP-phase in the neutrino oscillations directly probes the CP-phase in Leptogenesis!

# Summary

### Any physics behind?

 $\lambda_{1\mu} = \lambda_{2e} = 0$   $\ell_{L\mu} \quad \ell_{Le}$   $\ell_{L\tau} \quad N_{R1} \quad N_{R2} \quad \ell_{L\tau}$ 

A higher dimensional realization.

The charged leptons are on the branes.

The two right-handed neutrinos reside on the intersections.

The Higgs boson is not localized.

Once the observed  $\delta$  and m<sub>ee</sub> are found to be consistent with our predictions, they can be explained by the "surprisingly shaved" seesaw mechanism.

This might reflect the structure of spacetime geometry in higher dimensional theories...

# Backup

Sakharov three conditions ('67)

Density operator :  $\rho = \Sigma f_n | n > < n |$   $i \partial \rho / \partial t + [\rho, H] = 0$  $\rho(t) = e^{iHt} \rho e^{-iHt}$ 

Baryon asymmetry :  $\langle n_B \rangle(t) = Tr[\rho(t) B]$  with  $\langle n_B \rangle(0) = 0$ 

#### Sakharov three conditions ('67)

For 
$$[H, B] = 0$$
:  $< n_B > (t) = < n_B > (0) = 0$  Sakharov #1  
For  $[H, C] = 0$ :  $< n_B > (t) = - < n_B > (t) \rightarrow < n_B > (t) = 0$  Sakharov #2  
For  $[H, CP] = 0$ :  $< n_B > (t) = - < n_B > (t) \rightarrow < n_B > (t) = 0$ 

In thermal equilibrium : Baryon production rate = Inverse Baryon production rate

Sakharov #3

### Generic two-zero conditions

Normal Hierarchy	Inverted Hierarchy
$\lambda_{1\alpha} = 0$ $\tan z = \frac{\sqrt{m_2} U_{\alpha 2}^*}{\sqrt{m_3} U_{\alpha 3}^*} ,$	$\lambda_{1\alpha} = 0$ $\tan z = \frac{\sqrt{m_2} U_{\alpha 2}^*}{\sqrt{m_1} U_{\alpha 1}^*} ,$
$\lambda_{2\alpha} = 0$	$\lambda_{2\alpha} = 0$
$\tan z = -\frac{\sqrt{m_3}  U_{\alpha 3}^*}{\sqrt{m_2}  U_{\alpha 2}^*}$	$\tan z = -\frac{\sqrt{m_1}  U_{\alpha 1}^*}{\sqrt{m_2}  U_{\alpha 2}^*}$
$\lambda = \left(\begin{array}{ccc} a & a' & 0 \\ b & 0 & b' \end{array}\right)$	$\lambda = \left(\begin{array}{ccc} a & a' & 0 \\ b & 0 & b' \end{array}\right)$
$ \rightarrow m_2 U_{\alpha 2} U_{\alpha' 2} + m_3 U_{\alpha 3} U_{\alpha' 3} = 0 $	$ \rightarrow m_2 U_{\alpha 2} U_{\alpha' 2} + m_1 U_{\alpha 1} U_{\alpha' 1} = 0 $
$\lambda = \left(\begin{array}{ccc} a & 0 & 0 \\ b & b' & b'' \end{array}\right)$	$\lambda = \left(\begin{array}{ccc} a & 0 & 0 \\ b & b' & b'' \end{array}\right)$
$ \rightarrow \ U_{\alpha 2}  U_{\alpha' 3} = U_{\alpha 3}  U_{\alpha' 2} $	$ \rightarrow  U_{\alpha 2}  U_{\alpha' 1} = U_{\alpha 1}  U_{\alpha' 2} $

#### Definitions of the U<sub>MNS</sub>

$$U = \begin{bmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{bmatrix} \\ \times \operatorname{diag}(1, e^{i\frac{\alpha_{21}}{2}}, e^{\frac{\alpha_{31}}{2}}) .$$

$$\frac{|U_{e2}|^2}{|U_{e1}|^2} \equiv \tan^2 \theta_{12}; \quad \frac{|U_{\mu3}|^2}{|U_{\tau3}|^2} \equiv \tan^2 \theta_{23}; \quad U_{e3} \equiv \sin \theta_{13} e^{-i\delta},$$

Allowed Yukawa couplings  

$$\int_{|\delta/\pi|}^{4} \Delta c_{0.46,0.48,0.50,0.52,0.54,0.56,0.58,0.60}^{4} \text{, W} = 0$$

$$\int_{|\delta/\pi|}^{0} \int_{|\delta/\pi|}^{0} \int_{|\delta/\pi|}$$

In these cases, we have non-trivial very sharp predictions  $\delta \simeq \pm \pi/2 \qquad m_{ee} \simeq 47\,{
m meV}$ 

In the quark sector, the Cabbibo angle is a parameter.

The Cabbibo angle *can be derived* if we put zero in M<sub>d</sub> !

$$M_u = \begin{pmatrix} m_u & 0 \\ 0 & m_c \end{pmatrix} \qquad M_d = \begin{pmatrix} 0 & \sqrt{m_d m_s} \\ \sqrt{m_d m_s} & m_s \end{pmatrix}$$

→ 
$$sin\theta_{C} = (m_{d}/m_{s})^{1/2} \sim 0.22!$$

[S. Weinberg, HUTP-77-A057, Trans.New York Acad.Sci.38:185-201, 1977]

# Leptogenesis

$$\epsilon = \frac{\Gamma[N \to \ell + h] - \Gamma[N \to \ell^{\dagger} + h^{\dagger}]}{\Gamma[N \to \ell + h] + \Gamma[N \to \ell^{\dagger} + h^{\dagger}]}$$

$$\simeq \frac{3}{16\pi} \frac{M_1}{v^2} \frac{\mathrm{Im}[(\lambda m_{\nu} \lambda^T)_{11}]}{(\lambda \lambda^{\dagger})_{11}}$$

$$\frac{n_B}{n_{\gamma}} = \frac{28}{79} \frac{n_{B-L}}{n_{\gamma}} = \frac{28}{79} \frac{n_L}{n_{\gamma}} \Big|_{N_R \text{decay}}$$

 $P(v_{\mu} \rightarrow v_{e}) \approx sin^{2}2\theta_{13}T_{1} - \alpha sin2\theta_{13}T_{2} + \alpha sin2\theta_{13}T_{3} + \alpha^{2}T_{4}$ 

 $T_1 = sin^2 \theta_{23} sin^2 [(1-x_v)\Delta]/(1-x_v)^2$ 

 $T_2 = sin\delta sin2\theta_{12} sin2\theta_{23} sin\Delta sin(x_v\Delta)/x_v sin[(1-x_v)\Delta]/(1-x_v)$ 

 $T_3 = \cos\delta \sin 2\theta_{12} \sin 2\theta_{23} \cos\Delta \sin(x_{\nu}\Delta)/x_{\nu} \sin[(1-x_{\nu})\Delta]/(1-x_{\nu})$ 

 $T_4 = \cos^2\theta_{23} \sin^2 2\theta_{12} \sin^2(x_v \Delta) / x_v^2$ 

 $\Delta \equiv \Delta m^2{}_{31}L/4E, \alpha \equiv \Delta m^2{}_{21}/\Delta m^2{}_{31} \sim 1/30, x_v \equiv 2\sqrt{2}G_F N_e E/\Delta m^2{}_{31}$