

PLAN

- Introduction to $2\beta 0\nu$
- NEMO3
 - description, performances
 - results $2\beta 2\nu$
 - results $2\beta 0\nu$: data phase 1 1.08 year
 - fight against radon: result
- Prospects: SuperNEMO R&D
- Conclusion

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Why looking for $2\beta 0\nu$

1. Proton decay experiments

Super Kamiokande experiment

$p \not\rightarrow e^+\pi^0$ at a level of $1.6 \cdot 10^{33}$ years ($5.4 \cdot 10^{33}$ y updated)

$p \not\rightarrow \nu K^+$ at a level of $6.7 \cdot 10^{32}$ years

robust consequence SU(5) is not the right GUT

2. Neutrino oscillation experiments

ν_μ beam from accelerator K2K (≈ 1 GeV)

ν_e, ν_μ atmospheric K and S.K (\approx GeV)

ν_e from sun, reactors, Davis, Sage, Gallex, S. Kamioka, Chooz, SNO, Kamland (\approx MeV)

$\Rightarrow m_\nu \neq 0$ but a tiny mass

next soon OPERA $\nu_\mu \rightarrow \nu_\tau$?

miniboone ν_s ?

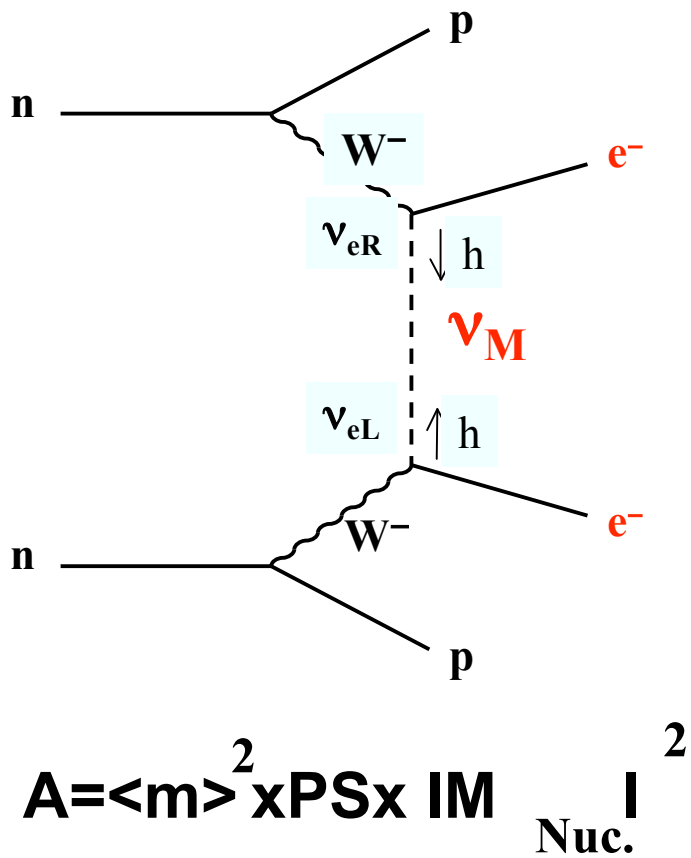
MINOS ν_μ disappearance

3. If GUT is the line guide then beyond SU5 ν 's are ν_M with a tiny mass

$\Rightarrow 2\beta 0\nu$ mass mechanism ?

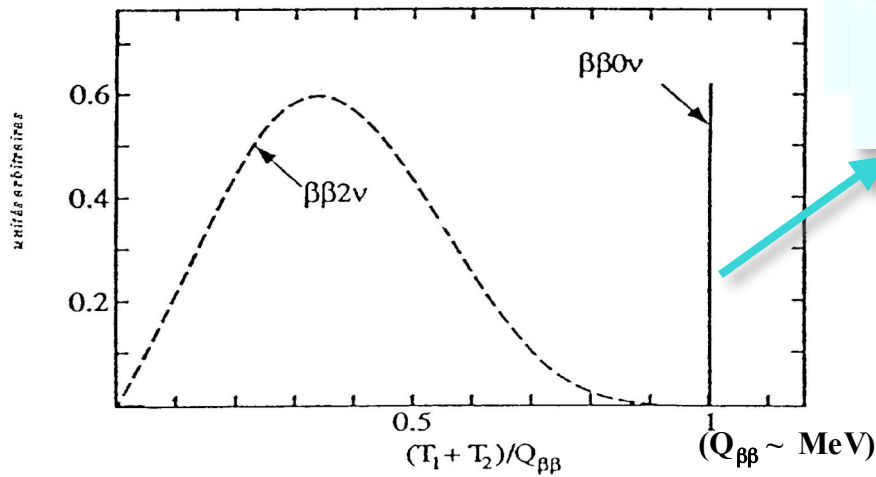
Double beta $\beta\beta(0\nu)$ decay: Physics beyond the standard model

$\beta\beta(0\nu) : 2n \rightarrow 2p+2e^-$



$\Delta L = 2$ Process

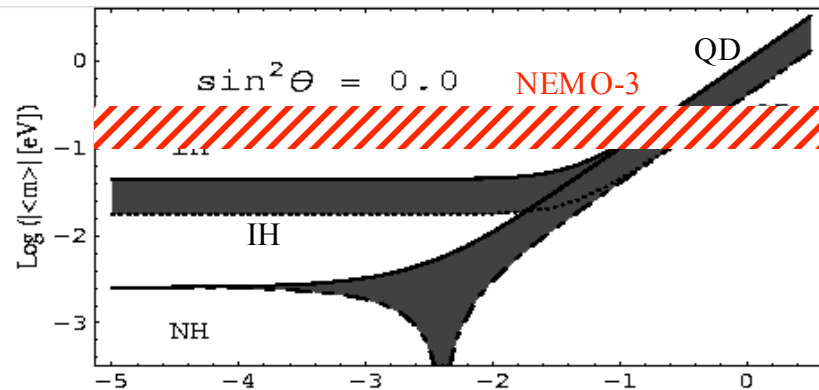
- Majorana Neutrino $\nu = \bar{\nu}$ and effective mass $\langle m_\nu \rangle$
- Right-handed current in weak interaction
- Majoron emission
- SUSY particle exchange



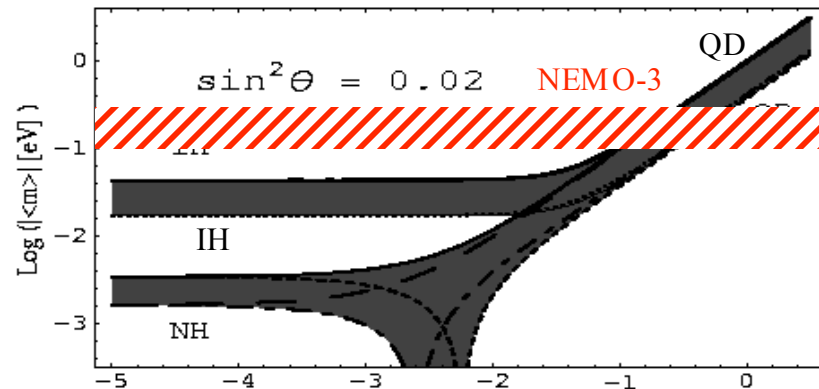
$(T_{1/2}^{0\nu})$

Expected values of $\langle m_\nu \rangle$ from neutrinos oscillations parameters

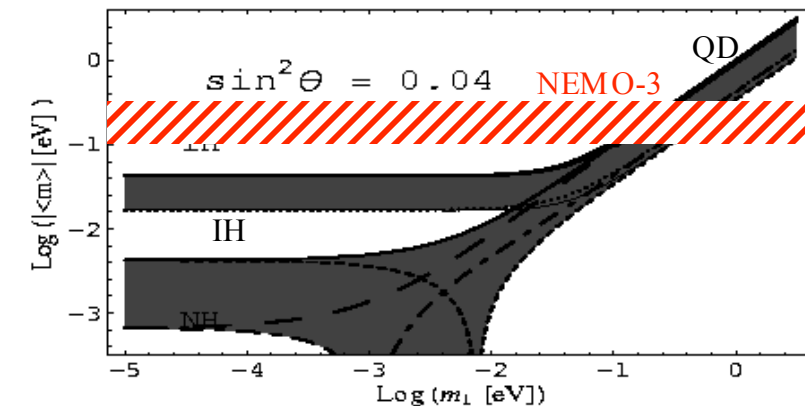
$\sin^2\theta_{13} = 0.0$



$\sin^2\theta_{13} = 0.02$



$\sin^2\theta_{13} = 0.04$



$$\langle m_\nu \rangle = \sum_1^3 m_i U_{ei}^2 e^{i\alpha_j}$$

Pascoli and Petcov, hep-ph/0310003
(best fit $\nu_{\text{atm}} + \nu_{\text{sol}}$)

Quasi-Degenerate (QD):

$$\langle m_\nu \rangle > 50 \text{ meV}$$

Inverted Hierarchy (IH):

$$15 \text{ meV} < \langle m_\nu \rangle < 50 \text{ meV}$$

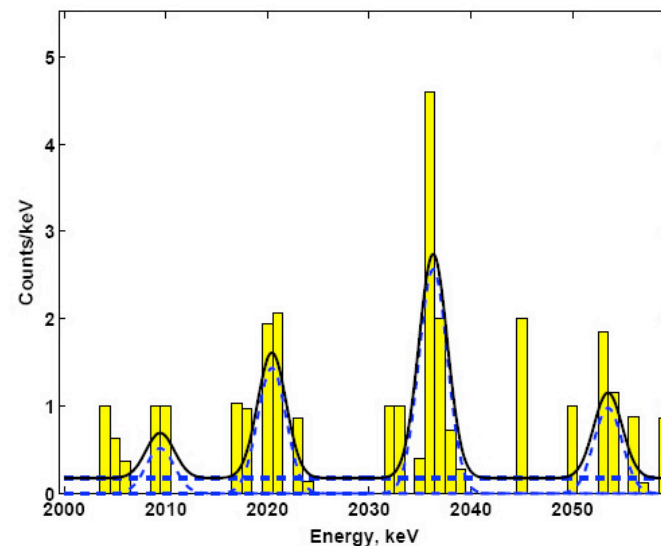
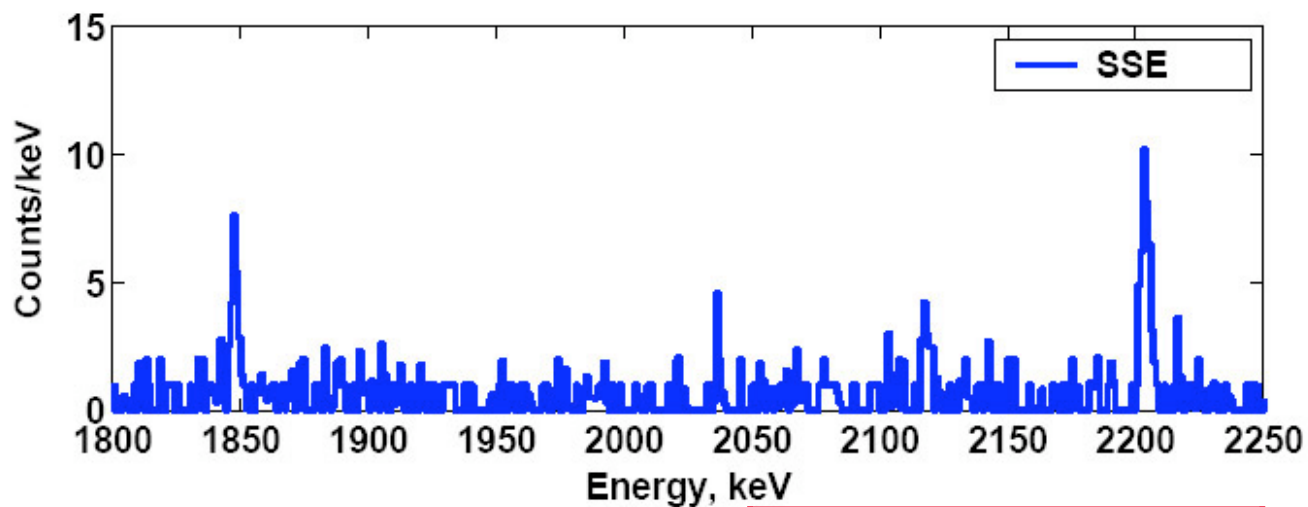
Normal Hierarchy (NH):

$$\langle m_\nu \rangle < 5 \text{ meV}$$

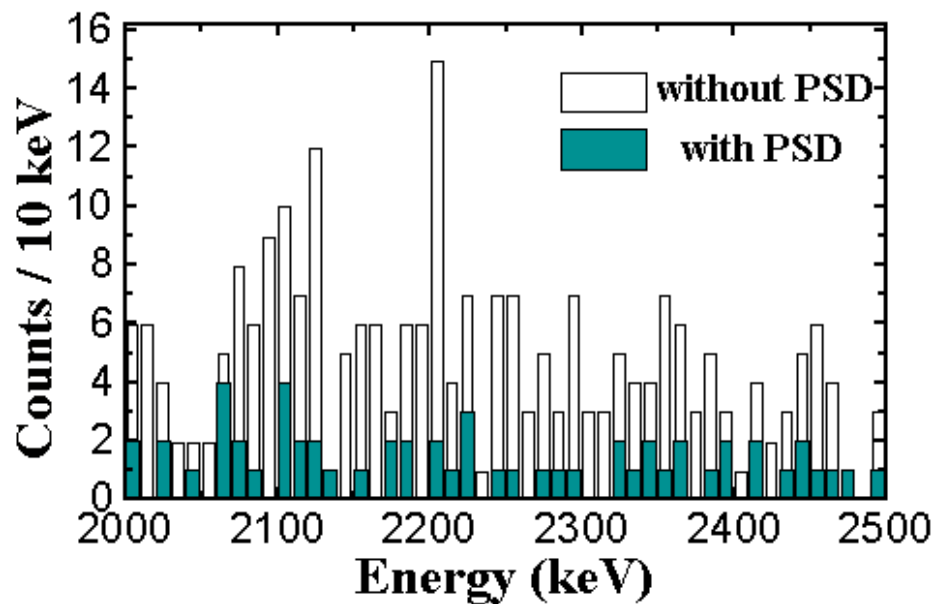
2β could give the absolute
neutrino mass

Present limits for different isotopes: direct measurements

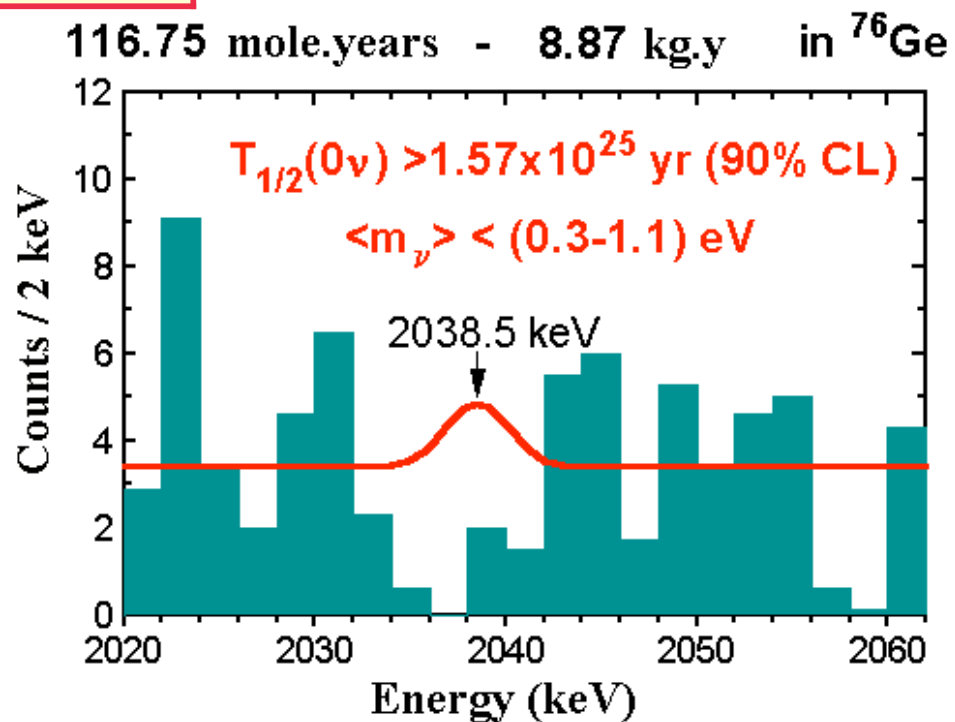
Isotope	$T_{1/2}(90\%CL)(y)$	$\langle m_\nu \rangle (eV)$	Mass (kg.y)	Experiment
^{48}Ca	$> 1.8 \cdot 10^{22}$	$< 6.3 - 39.4$	0.005	Candle
^{76}Ge	$> 1.9 \cdot 10^{25}$	$< 0.35 - 1.05$	35.5	Heidelberg- Moscou
^{76}Ge	$> 1.57 \cdot 10^{25}$	$< 0.3 - 1.1$	8.9	IGEX
^{82}Se	$> 1.9 \cdot 10^{23}$	$< 1.3 - 3.6$	0.9	NEMO3
^{96}Zr	$> 1.0 \cdot 10^{21}$	< 2.3	0.008	NEMO2
^{100}Mo	$> 3.5 \cdot 10^{23}$	$< 0.7 - 1.2$	7.5	NEMO3
^{116}Cd	$> 1.3 \cdot 10^{23}$	< 1.7	0.16	Solotvina
^{130}Te	$> 1.8 \cdot 10^{24}$	$< 0.2 - 1.1$	10.8	CUORICINO
^{136}Xe	$> 4.4 \cdot 10^{23}$	$< 2 - 3$	2.3	Gotthard
^{150}Nd	$> 1.2 \cdot 10^{21}$	< 3	0.009	TPC M.Moe



IGEX experiment



116.75 mole.years - 8.87 kg.y in ^{76}Ge



116.75 mole.years - 8.87 kg.y in ^{76}Ge

Running experiments

For mass ≈ 10 kg

Experiment	isotope	Mass (kg)	Type of detector	Lab.	Expected background (counts/ FWHM .kg.y)	Sensitivity $T_{1/2}$ (y)	Limit $\langle m_\nu \rangle$ (eV)
CUORICINO	^{130}Te	11	Bolometer	Gran Sasso (Italie)	0.5	7×10^{24}	0.3
NEMO3	^{100}Mo	6.9	Tracko-Calo	Modane (France)	0.3	4×10^{24}	0.20-0.35
	^{82}Se	0.93			0.1	8×10^{23}	0.65-1.8

NEMO3 Collaboration

CENBG, IN2P3-CNRS Bordeaux University, France

Charles University, Praha, Czech Republic

CTU, Praha, Czech Republic

INEL, Idaho Falls, USA

INR, Moscow, Russia

IReS, IN2P3-CNRS Strasbourg University, France

ITEP, Moscou, Russia

JINR, Dubna, Russia

Jyvaskyla University, Finland

LAL, IN2P3-CNRS Paris-Sud University, France

LSCE, CNRS Gif sur Yvette, France

LPC, IN2P3-CNRS Caen University, France

Manchester University, Great-Britain

Mount Holyoke College, USA

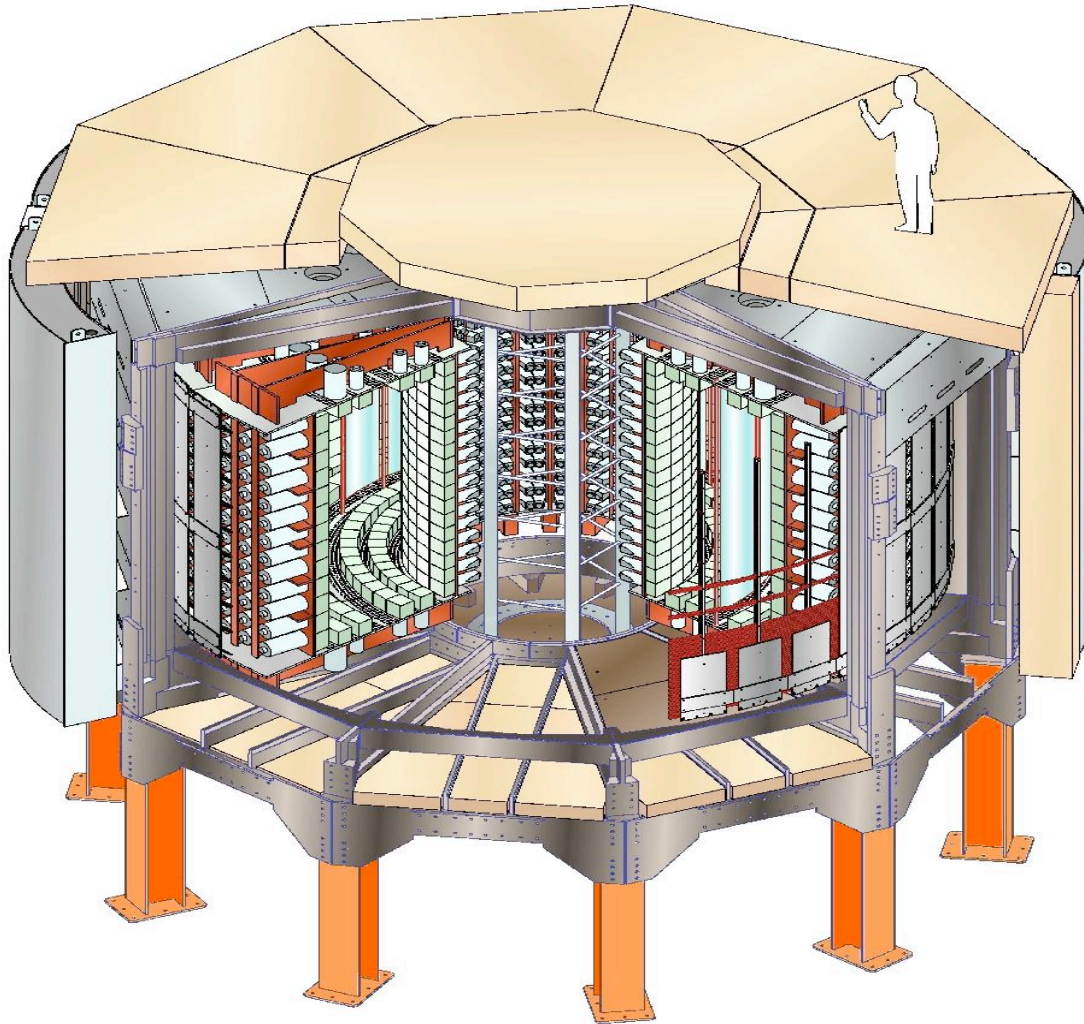
RRC kurchatov Institute, Moscow, Russia

Saga university, Saga, Japon

UCL, London, Great-Britain

The NEMO3 detector

Fréjus Underground Laboratory : 4800 m.w.e.



Source: 10 kg of $\beta\beta$ isotopes
cylindrical, $S = 20 \text{ m}^2$, $e \sim 60 \text{ mg/cm}^2$

Tracking detector:

drift wire chamber operating
in Geiger mode (6180 cells)
Gas: He + 4% ethyl alcohol + 1% Ar + 0.1% H₂O

Calorimeter:

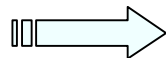
1940 plastic scintillators
coupled to low radioactivity PMTs

Magnetic field: 25 Gauss

Gamma shield: Pure Iron ($e = 18 \text{ cm}$)

Neutron shield: 30 cm water (ext. wall)

40 cm wood (top and bottom)
(since march 2004: water + boron)



Able to identify e^- , e^+ , γ and α



PMTs

scintillators

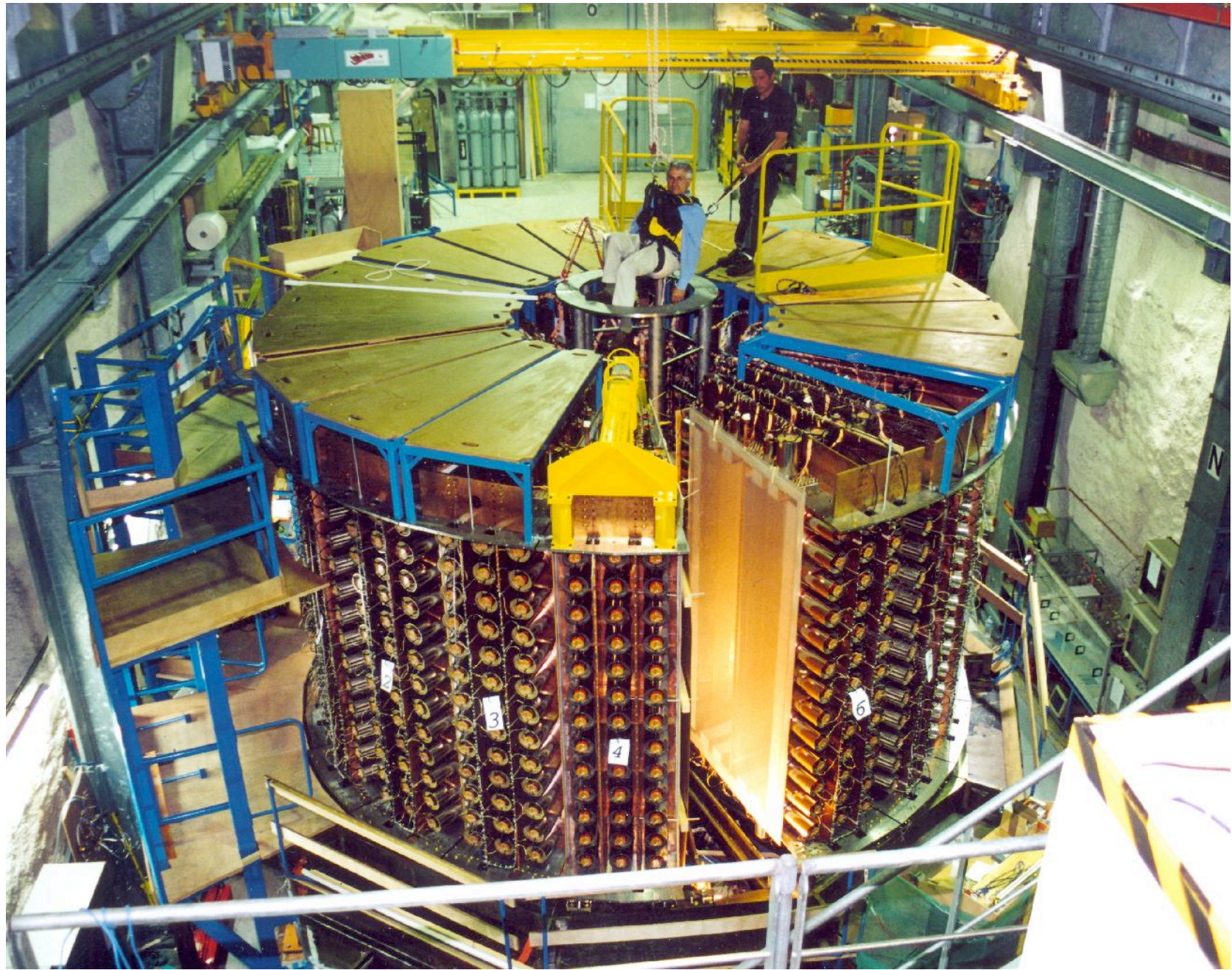
$\beta\beta$ isotope foils

**Cathodic rings
Wire chamber**

Calibration tube

Sources preparation

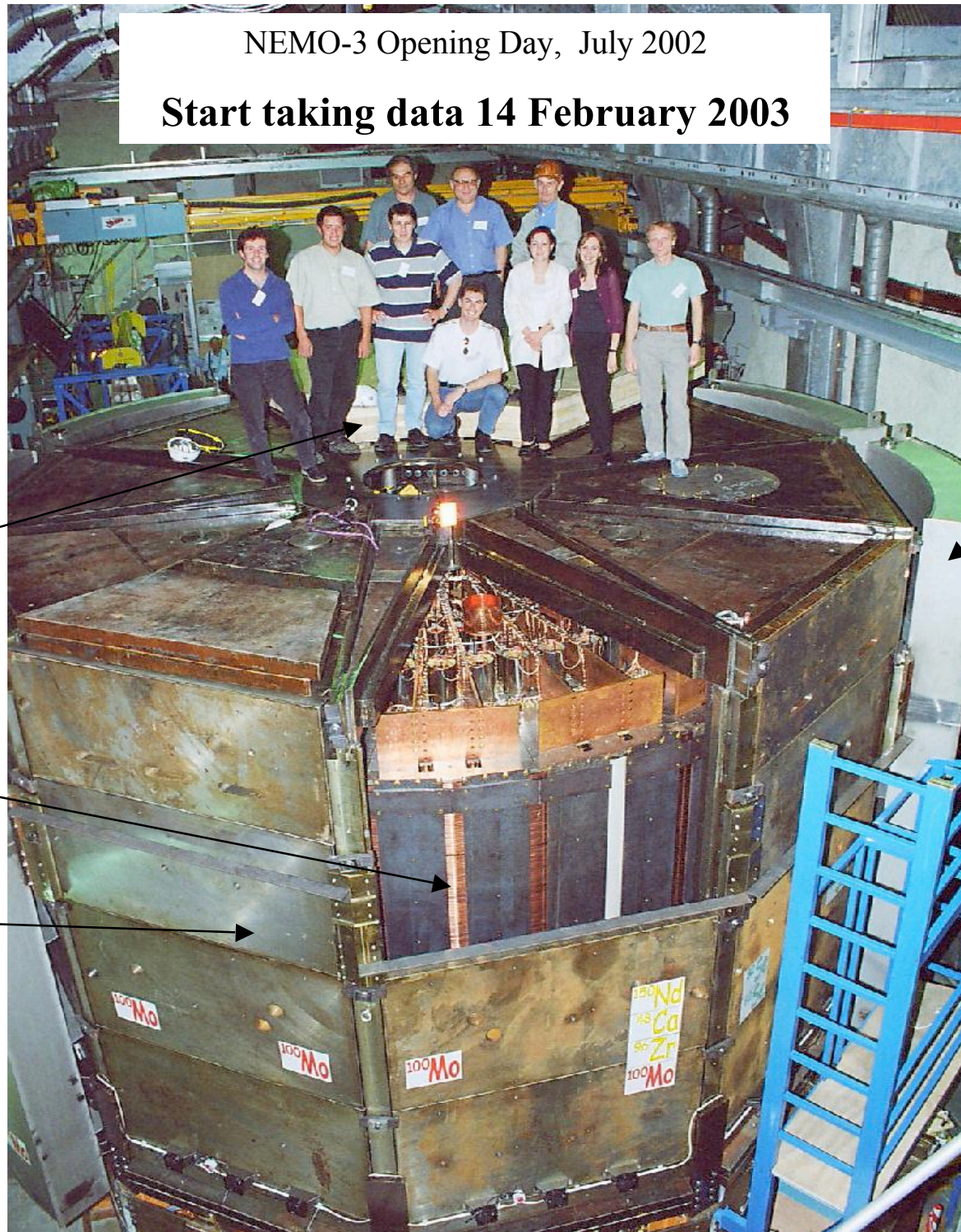




AUGUST 2001

NEMO-3 Opening Day, July 2002

Start taking data 14 February 2003



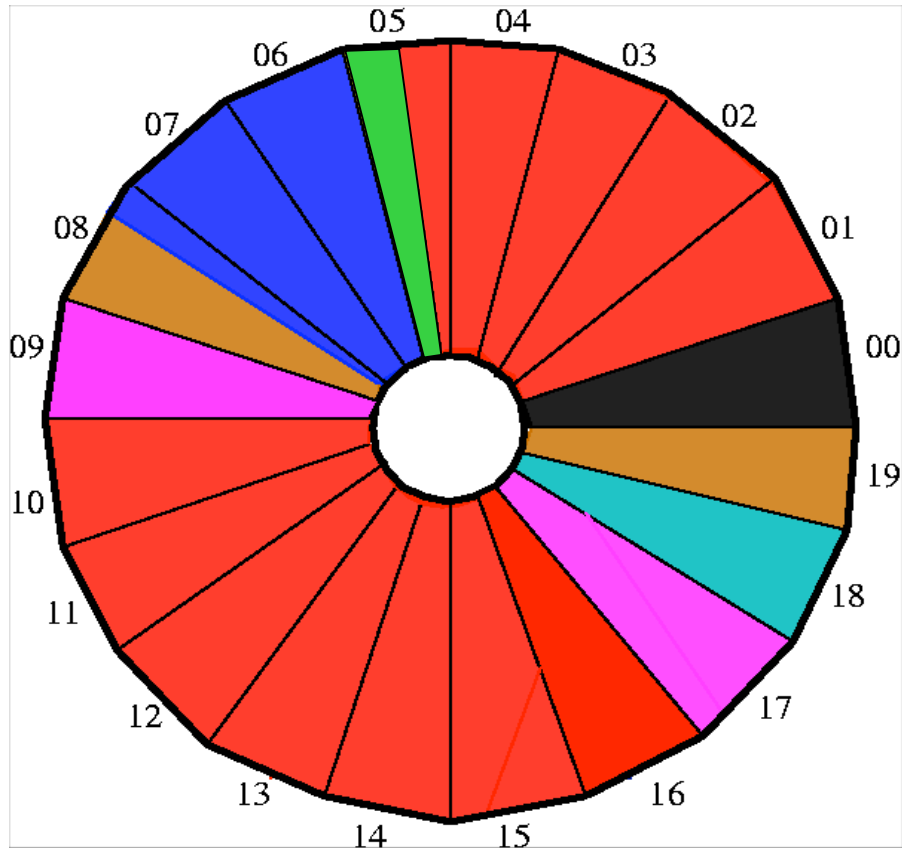
wood

coil

Iron shield

Water tank

ββ decay isotopes in NEMO-3 detector



^{100}Mo 6.914 kg **^{82}Se 0.932 kg**
 $Q_{\beta\beta} = 3034 \text{ keV}$ $Q_{\beta\beta} = 2995 \text{ keV}$

ββ0ν search

(All the enriched isotopes produced in Russia)

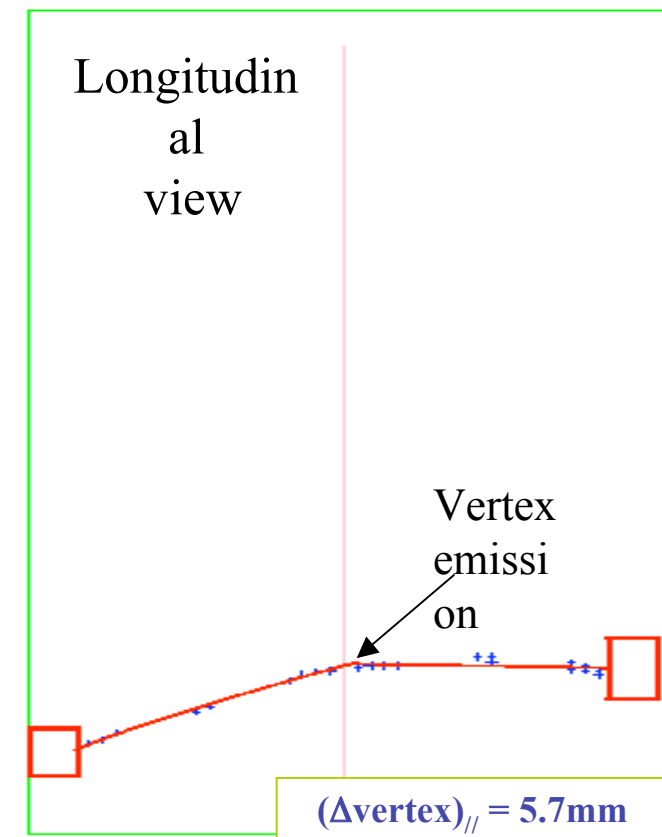
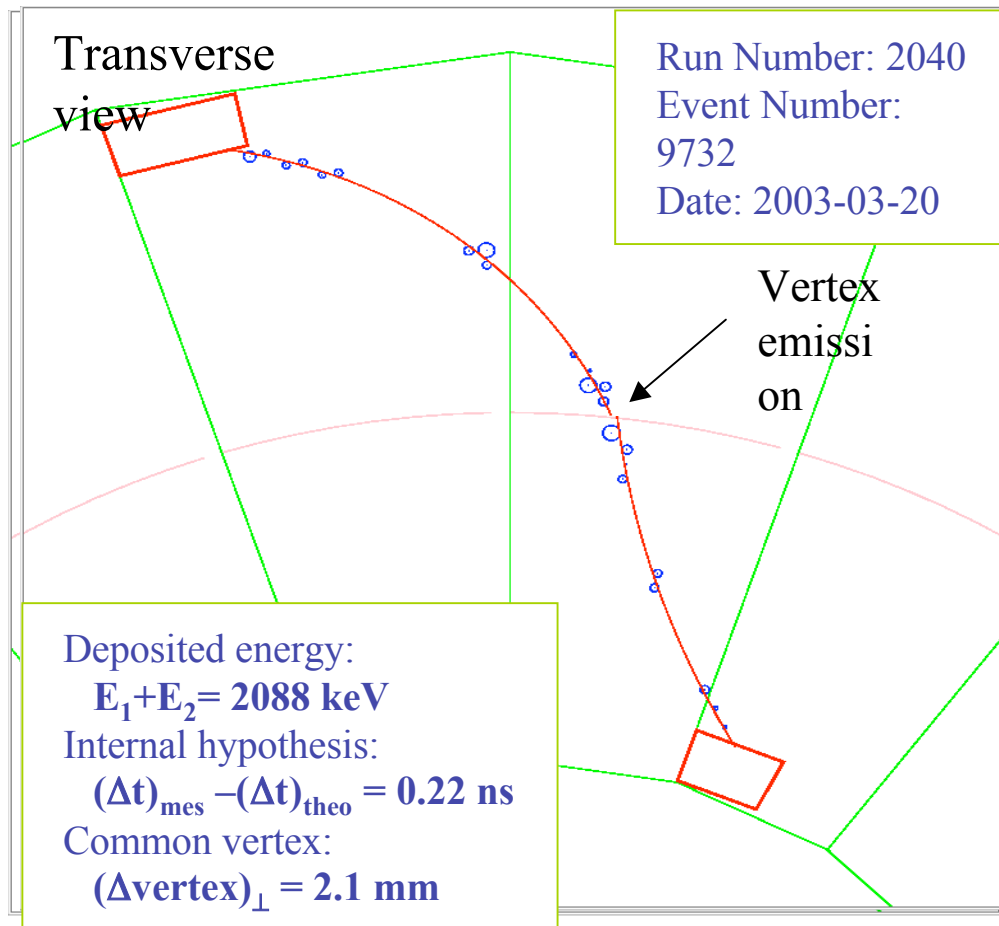
ββ2ν measurement

- ^{116}Cd 405 g**
 $Q_{\beta\beta} = 2805 \text{ keV}$
- ^{96}Zr 9.4 g**
 $Q_{\beta\beta} = 3350 \text{ keV}$
- ^{150}Nd 37.0 g**
 $Q_{\beta\beta} = 3367 \text{ keV}$
- ^{48}Ca 7.0 g**
 $Q_{\beta\beta} = 4272 \text{ keV}$
- ^{130}Te 454 g**
 $Q_{\beta\beta} = 2529 \text{ keV}$
- natTe 491 g**
- Cu 621 g**

External bkg measurement

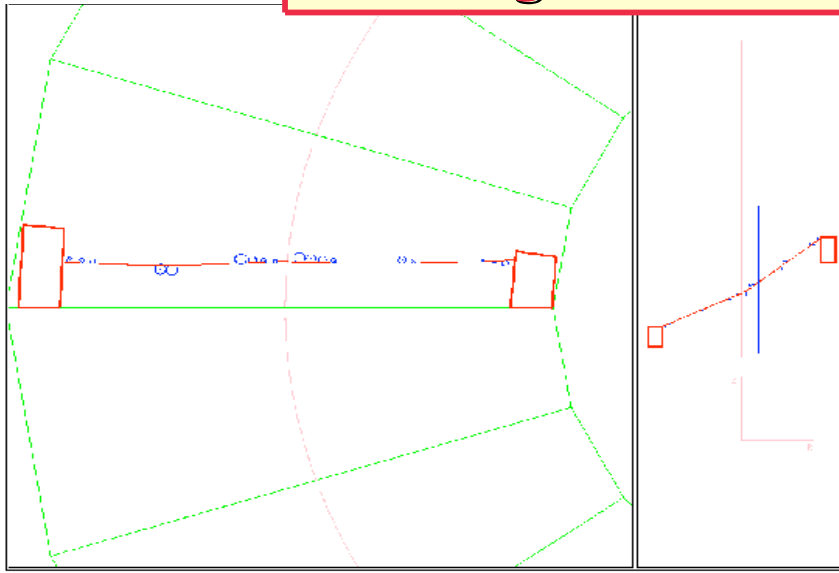
$\beta\beta$ events selection in NEMO-3

Typical $\beta\beta 2\nu$ event observed from ^{100}Mo

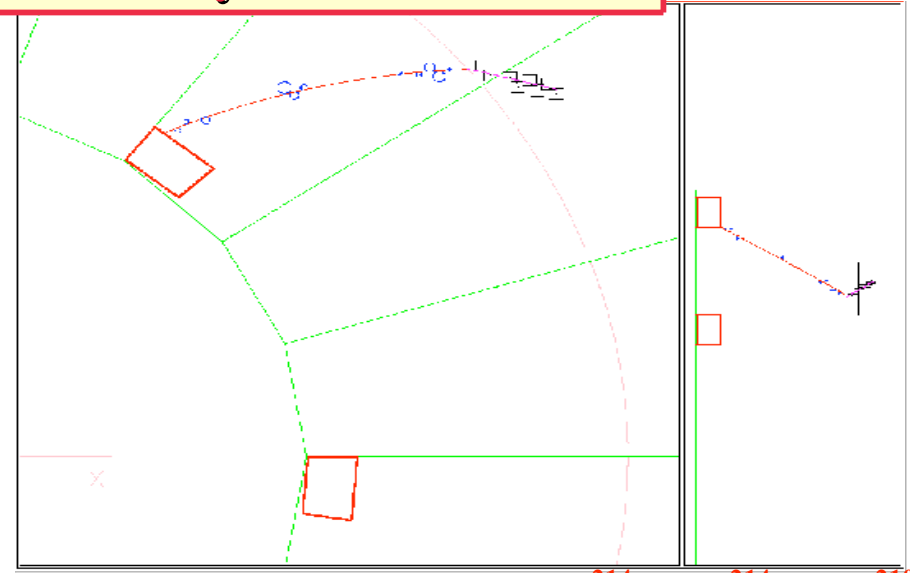


Trigger: 1 PMT > 150 keV
3 Geiger hits (2 neighbour layers + 1)
Trigger rate = 7 Hz
 $\beta\beta$ events: 1 event every 1.5 minutes

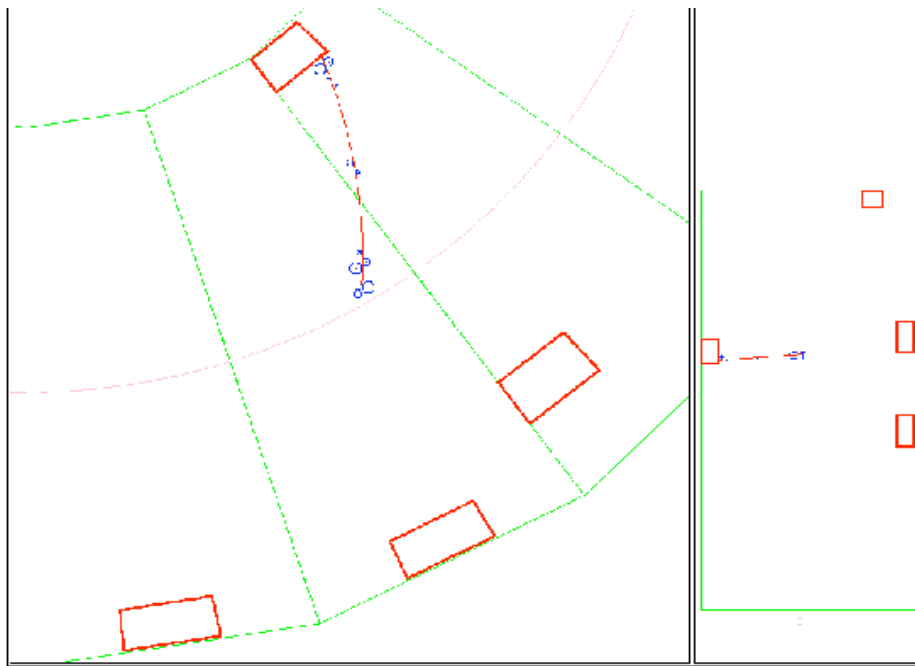
Background events observed by NEMO-3...



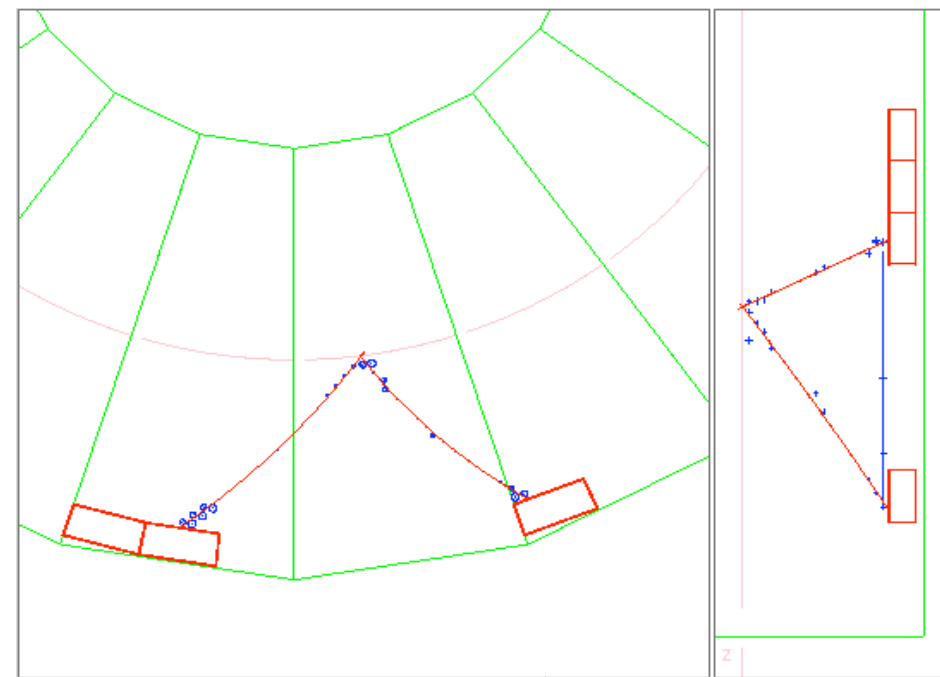
Electron crossing > 4 MeV **Neutron capture**



Electron + α delay track ($164 \mu\text{s}$) $^{214}\text{Bi} \rightarrow ^{214}\text{Po} \rightarrow ^{210}\text{Pb}$



Electron + N γ 's ^{208}Tl ($E_\gamma = 2.6$ MeV)



Electron - positron pair **B rejection**

NEMO3: detector performances RUN with radioactive sources

- energy calibration : absolute

^{207}Bi $2e^-$ lines $\approx 0.5 \approx 1$ MeV

^{90}Sr β^- end point ≈ 2.2 MeV

- time of flight

^{60}Co γ_1, γ_2 2 lines ≈ 1.5 MeV

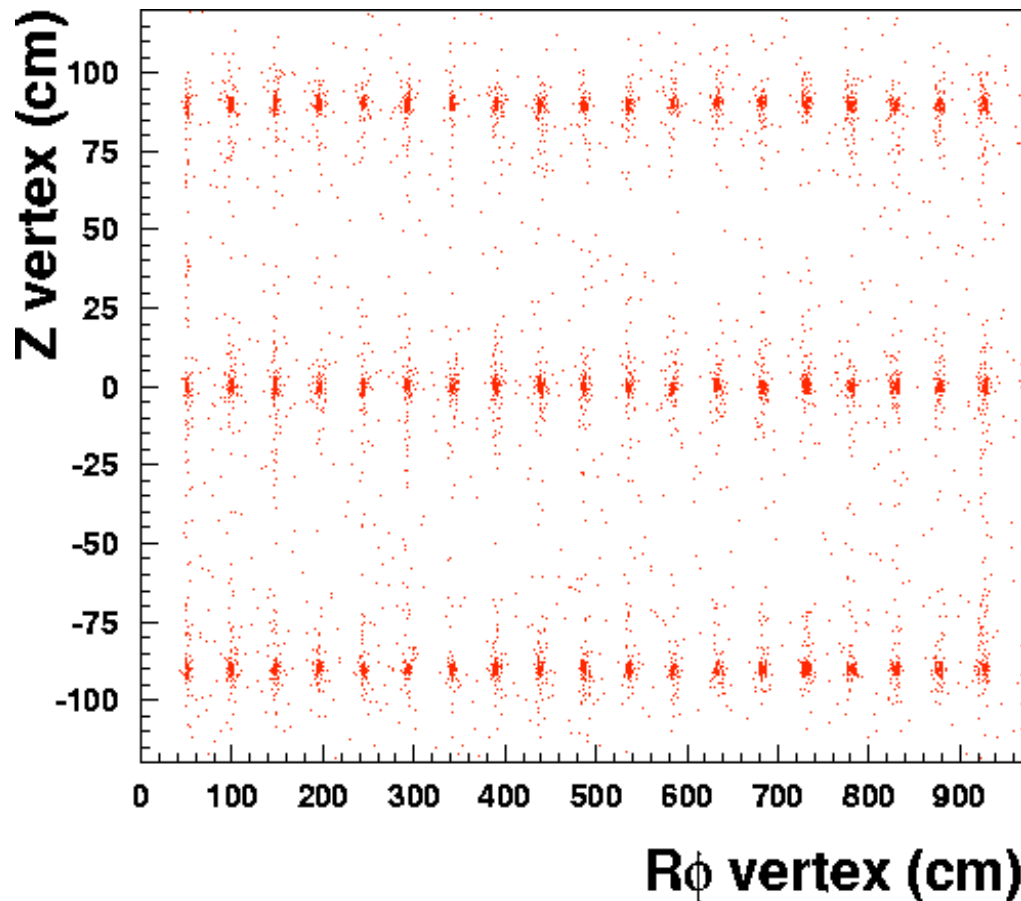
- tracking detector: $\sigma_T, \sigma_{//}$

neutron source \longrightarrow (n, γ) \longrightarrow $e^- > 4.5$ MeV

- aging? absolute calibration ^{207}Bi during 2 years

Transversal and Longitudinal Resolution on the Vertex

^{207}Bi sources at 3 well known positions in each sector
(emission of two e- conversion at ≈ 1 and 0.5 MeV)



1 e⁻ channel at 1 Mev:

$$\sigma_{\perp} (1 \text{ MeV}) = 0.25 \text{ cm}$$

$$\sigma_{//} (1 \text{ MeV}) = 0.95 \text{ cm} \quad (Z=0)$$

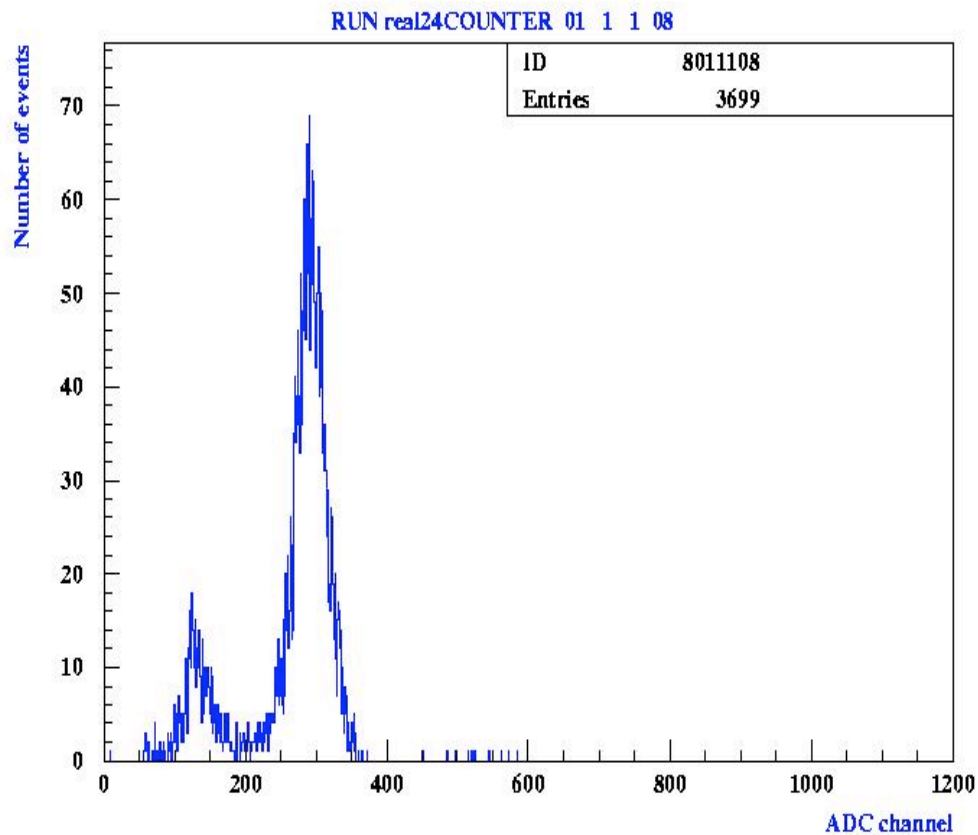
2e- channel (1 MeV+ 0.5 MeV)

$$\sigma_{\perp} (1 \text{ MeV}) = 0.6 \text{ cm}$$

$$\sigma_{//} (1 \text{ MeV}) = 1.3 \text{ cm} \quad (Z=0)$$

Performances of the calorimeter

Tube in each sector where calibration sources are introduced (3 positions)
3 electron energies : 486 keV and 976 keV with ^{207}Bi , and 2.28 MeV with ^{90}Sr

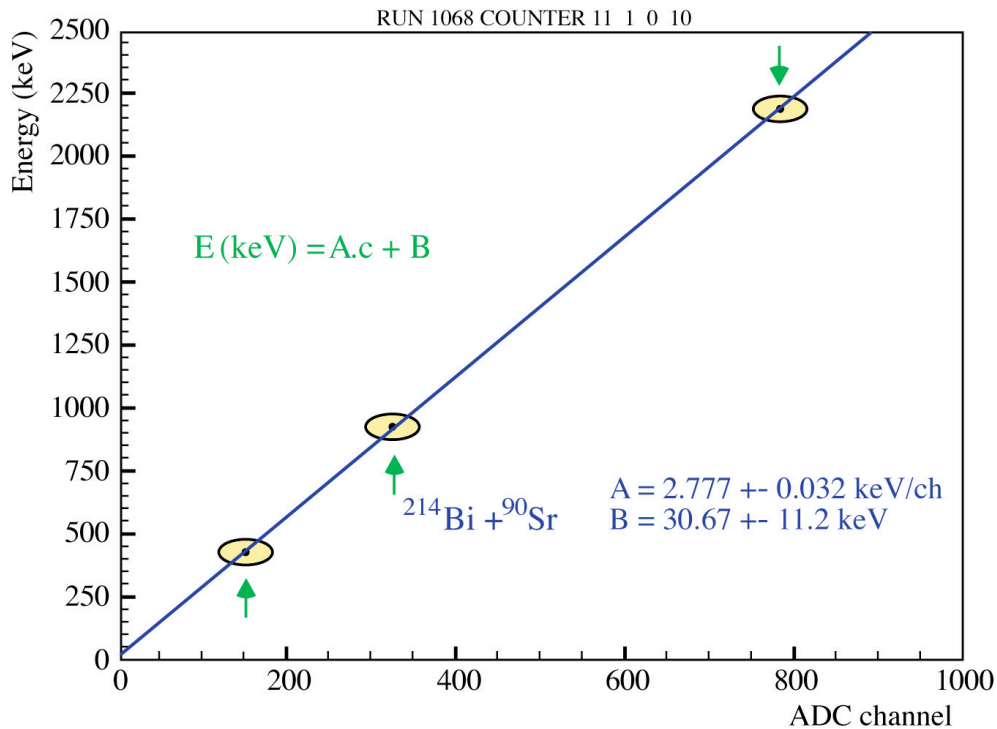
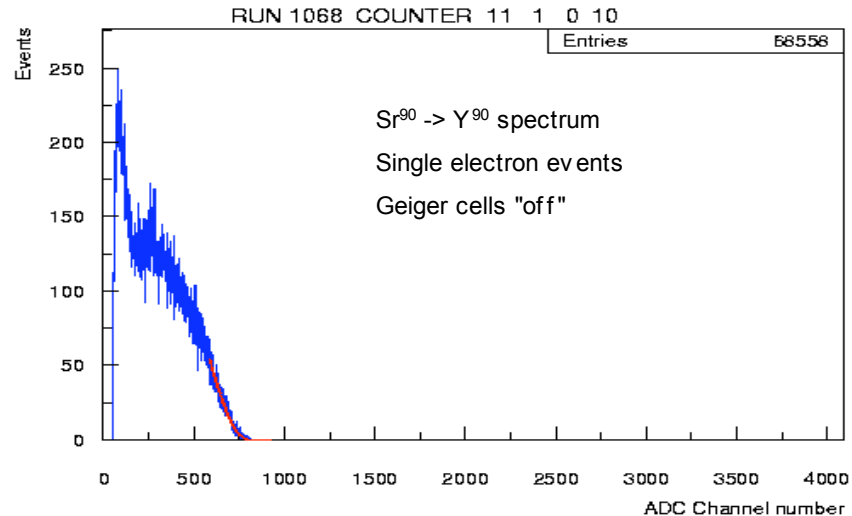
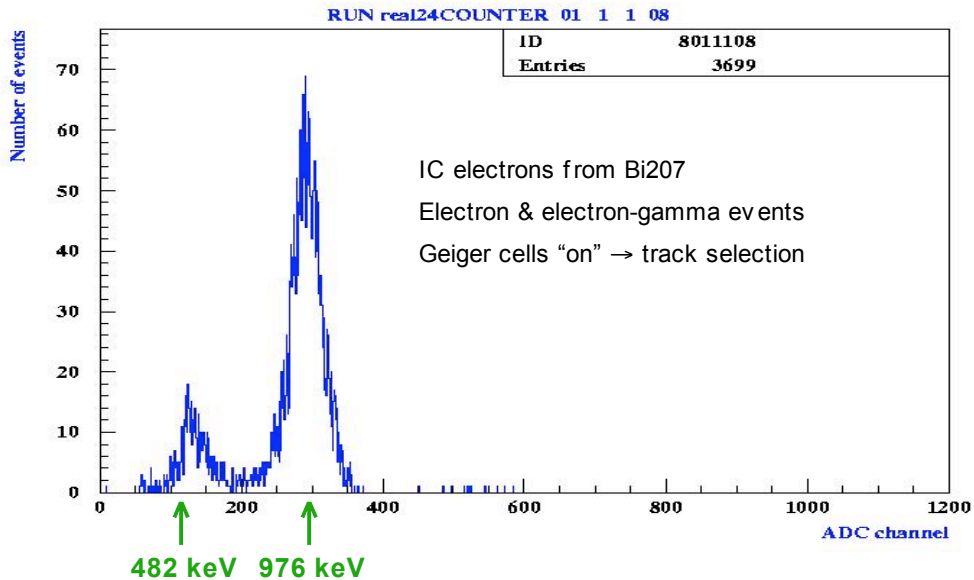


At 1 MeV ($Q_{\beta\beta} \approx 3$ MeV for ^{100}Mo and ^{82}Se):

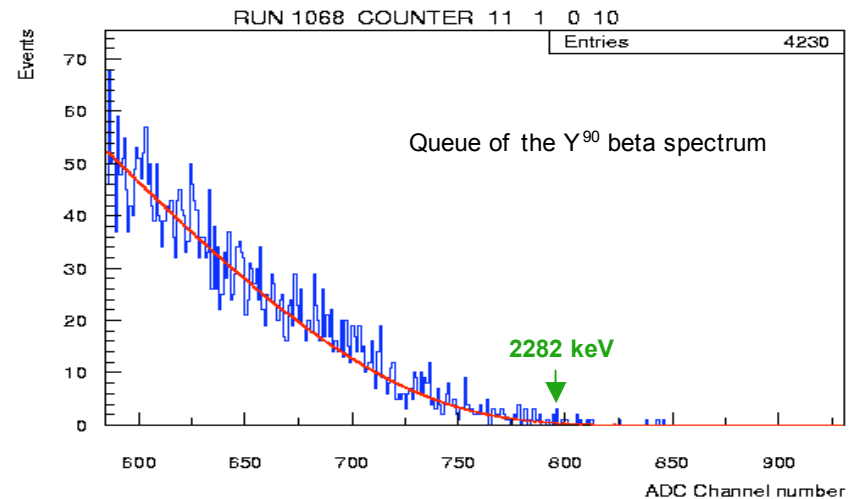
	FWHM	σ_E/E
Ext. Wall (PMTs 5")	14 %	$5.8 \% \sqrt{E(\text{MeV})}$
Int. Wall (PMTs 3")	17 %	$7.1 \% \sqrt{E(\text{MeV})}$

Absolute calibration: method

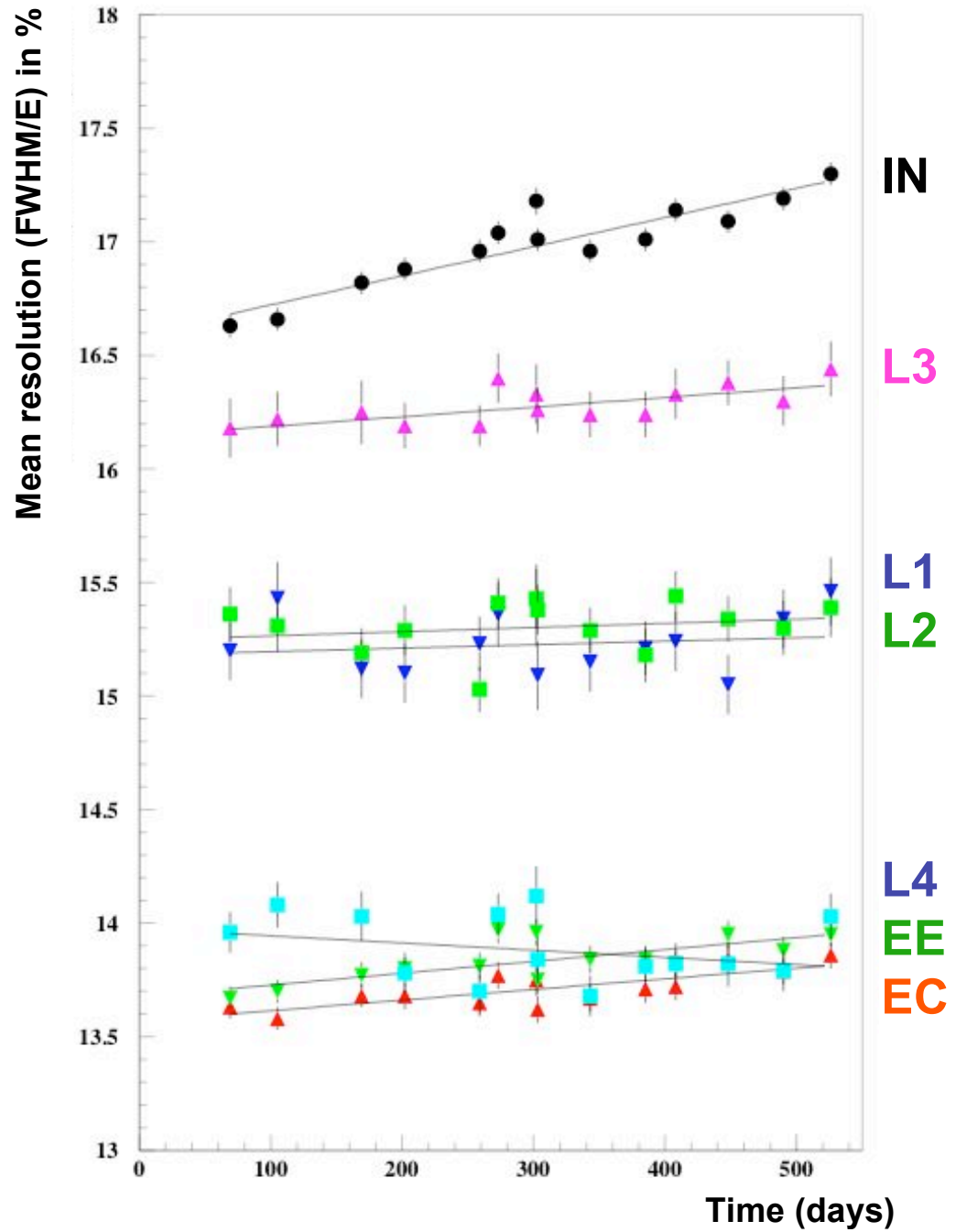
2003/09/08 11.51



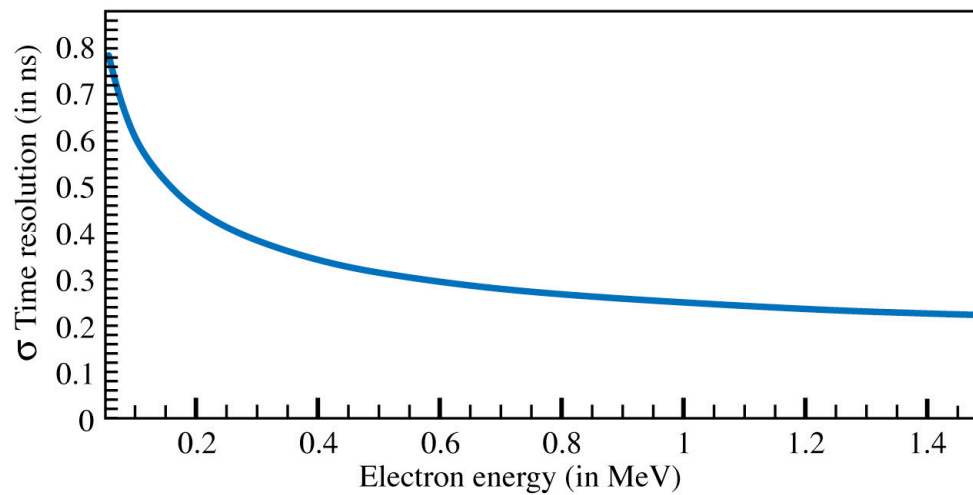
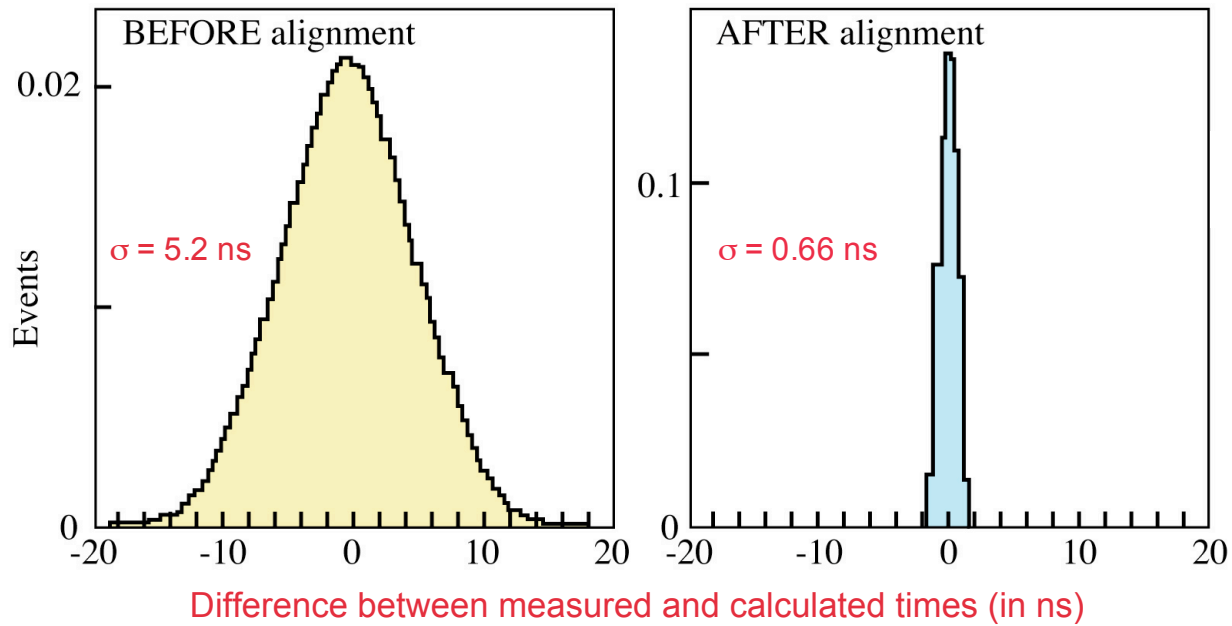
2003/09/08 11.51



Aging measurements



^{60}Co run: alignment of the 1940 units calo



Time resolution (in ns) as a function of the electron energy (in MeV), obtained with two-electron events

Performance of the detector

Tracking Detector:

➤ 99.5 % Geiger cells ON

➤ Vertex resolution:

2 e^- channels (482 and 976 keV) using ^{207}Bi sources
at 3 well known positions in each sector

$$\sigma_{\perp} (\Delta\text{Vertex}) = 0.6 \text{ cm}$$

$$\sigma_{\parallel} (\Delta\text{Vertex}) = 1.3 \text{ cm} \quad (Z=0)$$

➤ e^+/e^- separation with a magnetic field of 25 G
~ 3% confusion at 1 MeV

Calorimeter:

➤ 97% of the PMTs+scintillators are ON

➤ Energy Resolution:

calibration runs (every ~ 40 days) with ^{207}Bi
sources

	Ext. Wall 5" PMTs	Int. Wall 3" PMTs
FWHM (1 MeV)	14%	17%

➤ Daily Laser Survey to control gain stability of each PM

Time Of Flight:

➤ Time Resolution ($\beta\beta$ channel) \approx 250 ps at 1 MeV

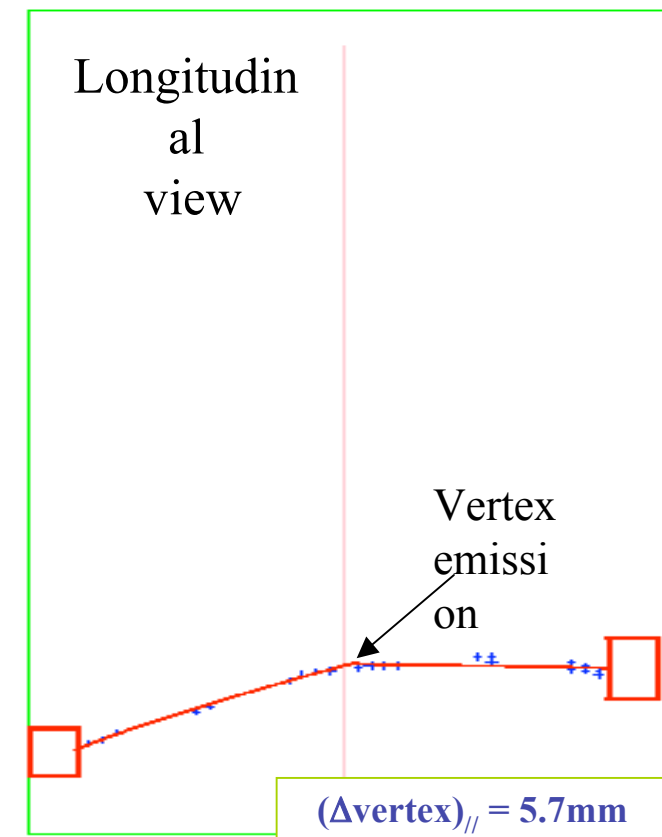
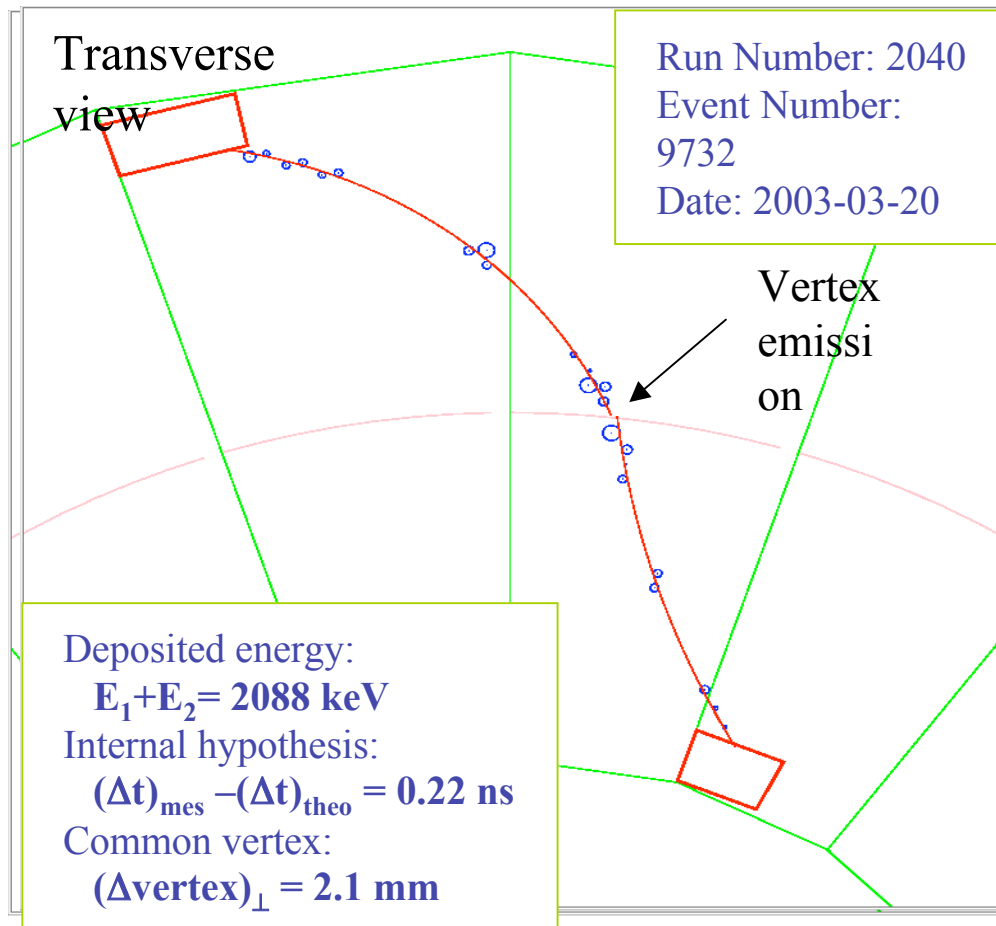
ToF (external crossing e^-) $>$ 3 ns

external crossing e^- totally rejected

**Expected Performance of the detector
has been reached**

$\beta\beta$ events selection in NEMO-3

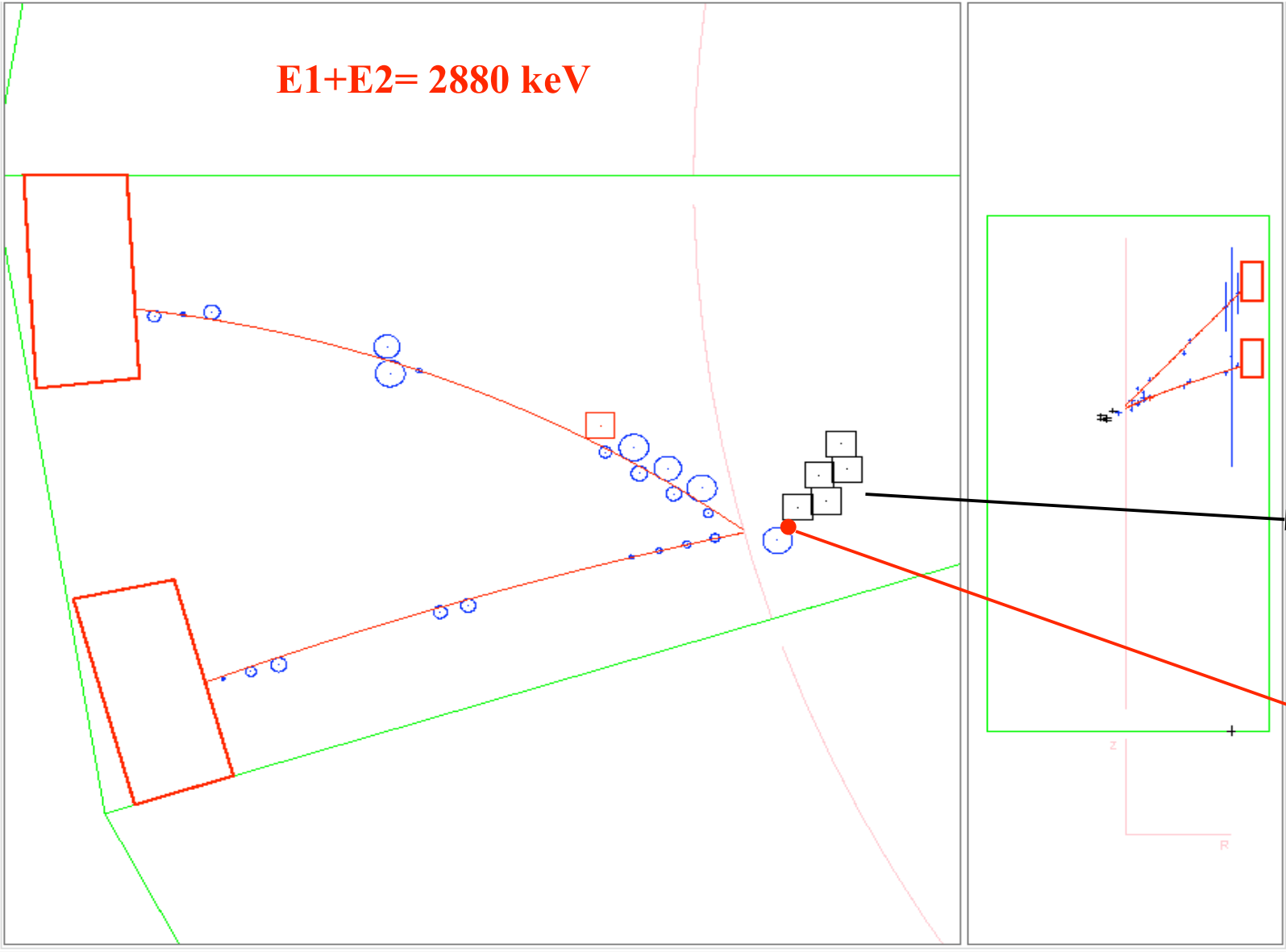
Typical $\beta\beta 2\nu$ event observed from ^{100}Mo



Trigger: 1 PMT > 150 keV
3 Geiger hits (2 neighbour layers + 1)
Trigger rate = 7 Hz
 $\beta\beta$ events: 1 event every 1.5 minutes

a $\beta\beta 0\nu$ -like event due to Radon from the gas

Run 2220, event 136.604, May 11th 2003



$E1+E2= 2880 \text{ keV}$

α track
(delay = $70 \mu\text{s}$)
 $^{214}\text{Po} \rightarrow ^{210}\text{Pb}$

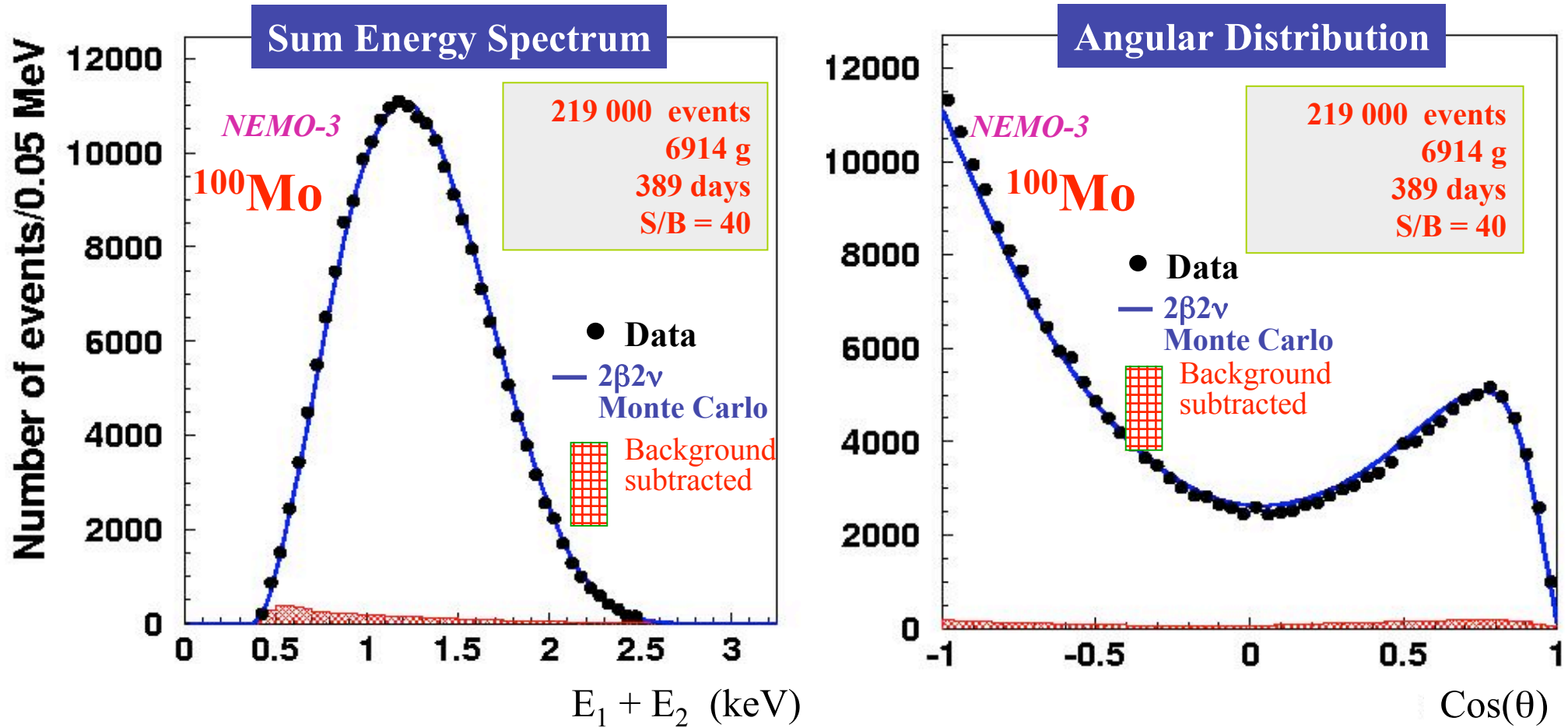
$^{214}\text{Bi} \rightarrow ^{214}\text{Po}$
 β decay
IN THE GAS

Event selection criteria

- Two tracks of negative charge associated to isolated PM
- Energy deposit in each scintillator $E > 200$ keV.
- Event vertex is inside the foil
- Distance track-to-vertex: $\Delta XY < 4$ cm, $\Delta Z < 8$ cm;
- TOF cut: internal hypothesis probability $> 4\%$, external hypothesis probability $< 1\%$;
- Reject events with the alpha particle found using *alpha_search* means:
 - if only 1 extra hit in the tracking detector $\left\{ \begin{array}{l} \Delta t > 40 \mu\text{sec} \\ \Delta xy < 4 \text{ cm} \\ \Delta Z < 10 \text{ cm} \end{array} \right\}$ vertex
 - if at least 2 hits
search for a short track
 $\Delta t > 2 \mu\text{sec}$ only but all hits on time
- Reject events with two tracks at one side of the foil and a geiger hit in time at the opposite side fo the foil close to the vertex: Möller scattering of β decay in gas (Radon).

^{100}Mo $2\beta 2\nu$ preliminary results

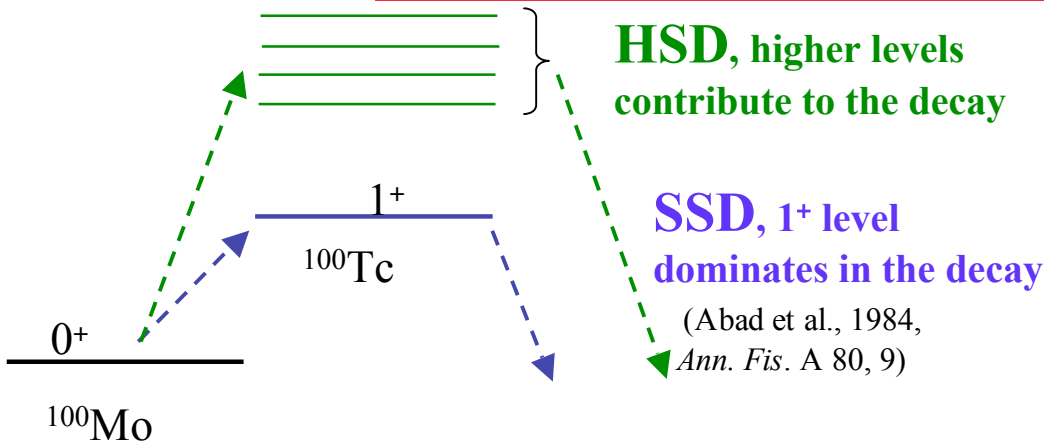
(Data Feb. 2003 – Dec. 2004)



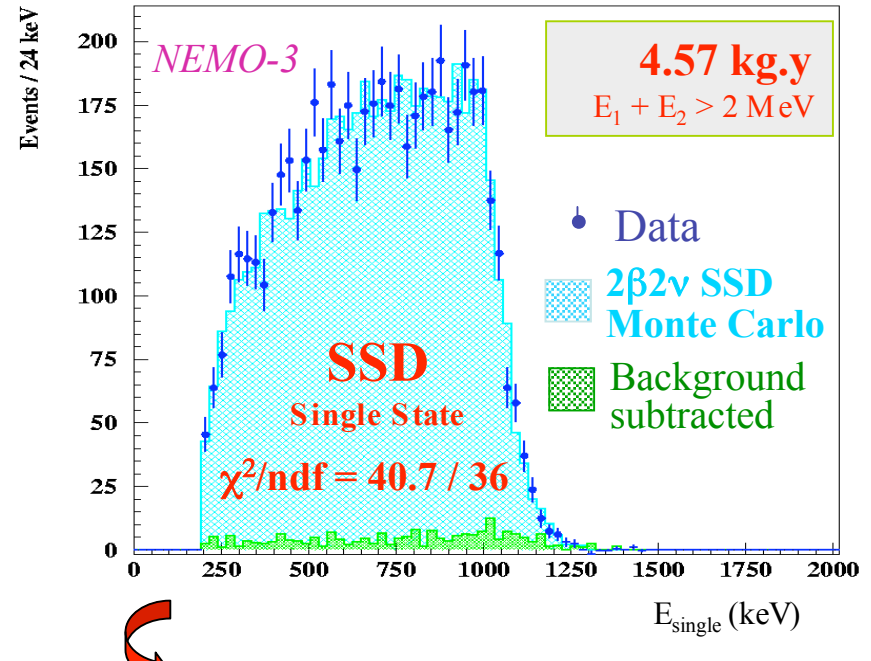
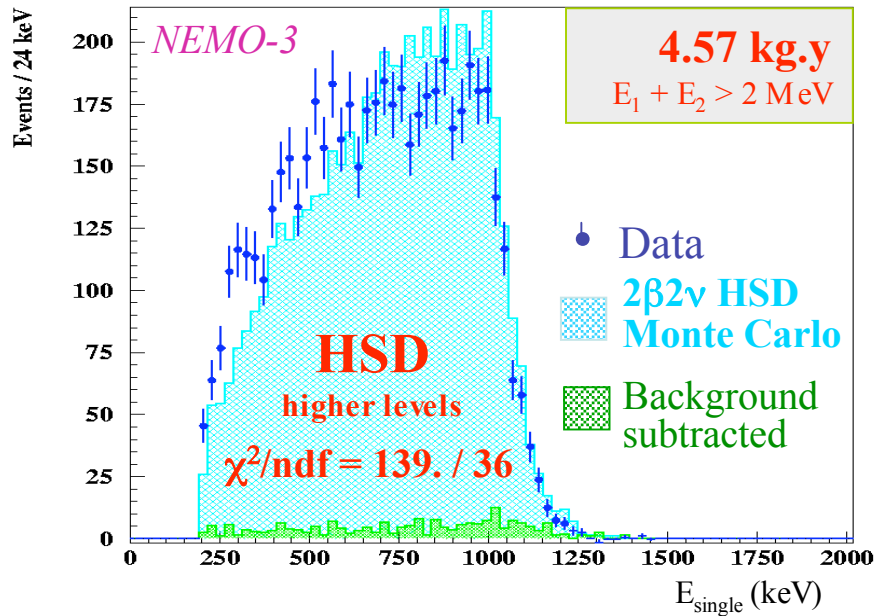
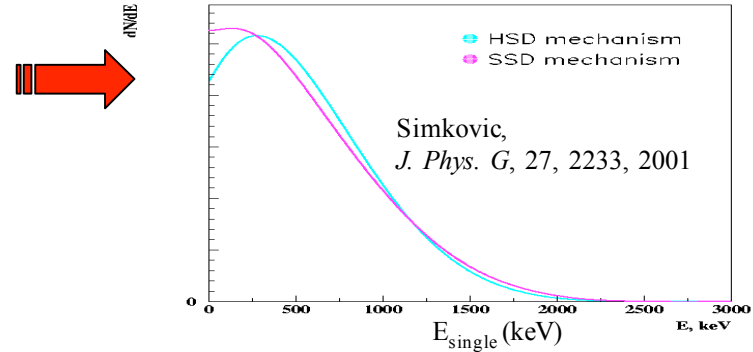
7.37 kg.y

$$T_{1/2} = 7.11 \pm 0.02 \text{ (stat)} \pm 0.54 \text{ (syst)} \times 10^{18} \text{ y}$$

^{100}Mo $2\beta 2\nu$ Single Energy Distribution



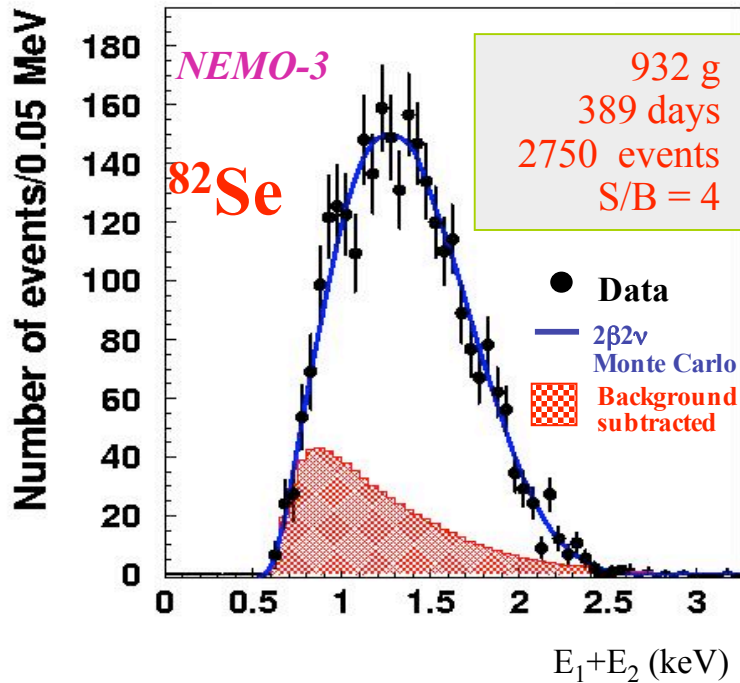
Single electron spectrum different between SSD and HSD



$\left\{ \begin{array}{l} \text{HSD: } T_{1/2} = 8.61 \pm 0.02 \text{ (stat)} \pm 0.60 \text{ (syst)} \times 10^{18} \text{ y} \\ \text{SSD: } T_{1/2} = 7.72 \pm 0.02 \text{ (stat)} \pm 0.54 \text{ (syst)} \times 10^{18} \text{ y} \end{array} \right.$

^{100}Mo $2\beta 2\nu$ single energy distribution
 in favour of Single State Dominant (SSD) decay

2 β 2 ν preliminary results for other nuclei



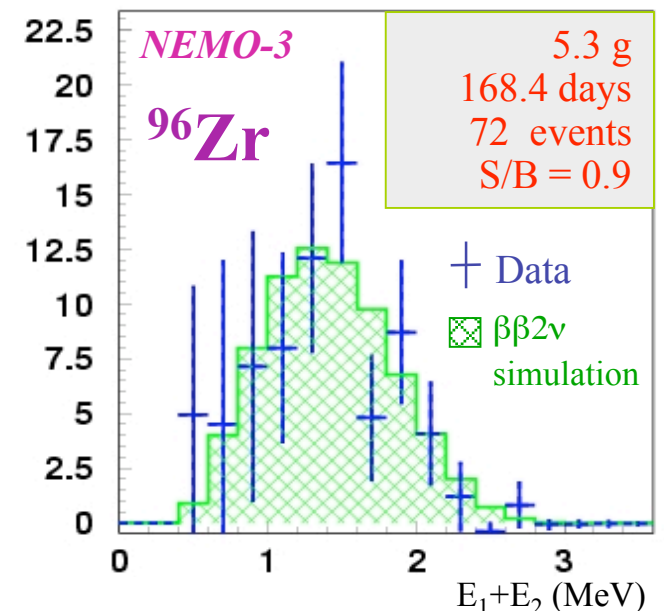
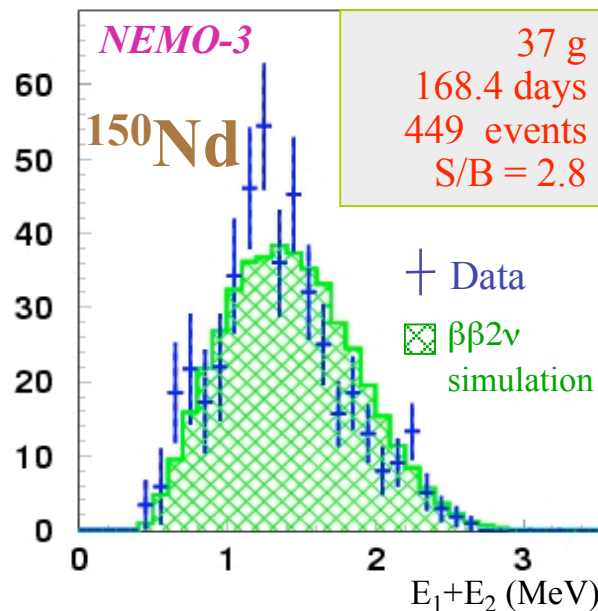
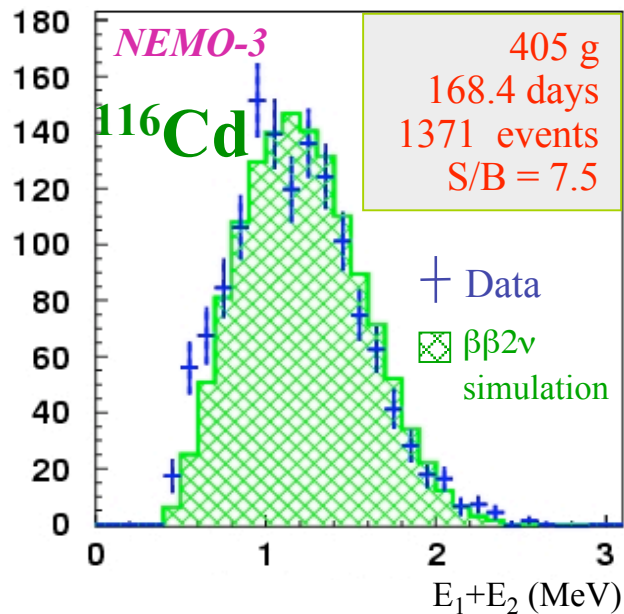
^{82}Se $T_{1/2} = 9.6 \pm 0.3$ (stat) ± 1.0 (syst) $\times 10^{19}$ y

^{116}Cd $T_{1/2} = 2.8 \pm 0.1$ (stat) ± 0.3 (syst) $\times 10^{19}$ y

^{150}Nd $T_{1/2} = 9.7 \pm 0.7$ (stat) ± 1.0 (syst) $\times 10^{18}$ y

^{96}Zr $T_{1/2} = 2.0 \pm 0.3$ (stat) ± 0.2 (syst) $\times 10^{19}$ y

Background subtracted



Ca48 analysis 1st preliminary result

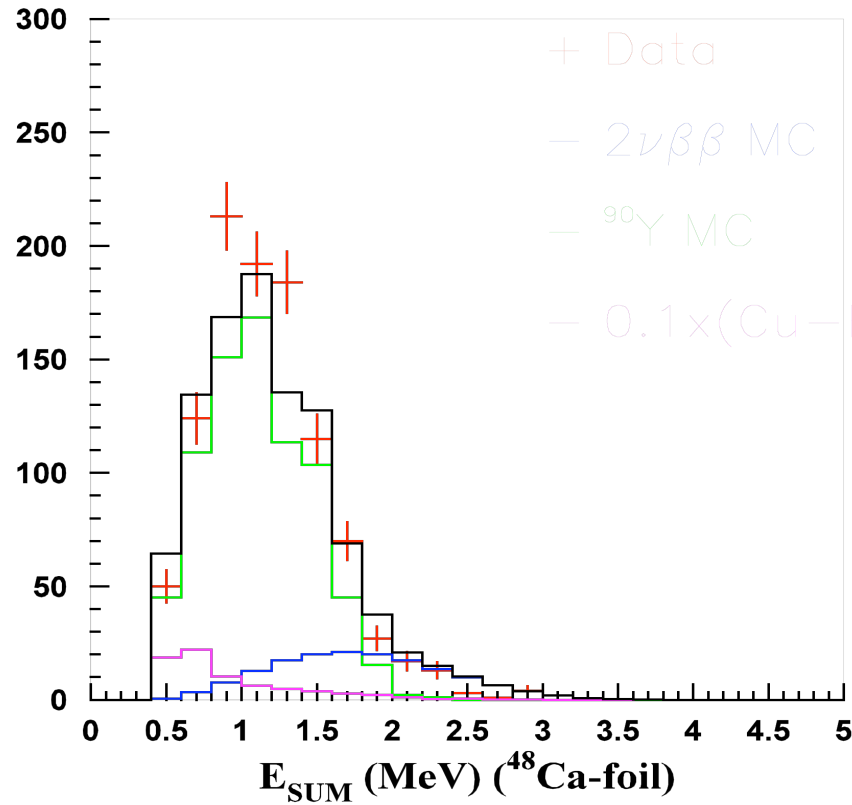
Hideaki Ohsumi analysis

- why $2\nu\beta 2\nu$ ^{48}Ca : test of the shell model calculation
- 7g of ^{48}Ca enough radiopure after chemistry ^{214}Bi , ^{208}Tl but 30m Bq of ^{90}Sr ! pure beta emitter criteria to remove Möller scattering E1 and E1 > 0.7 MeV $\cos\theta < 0$ back to back

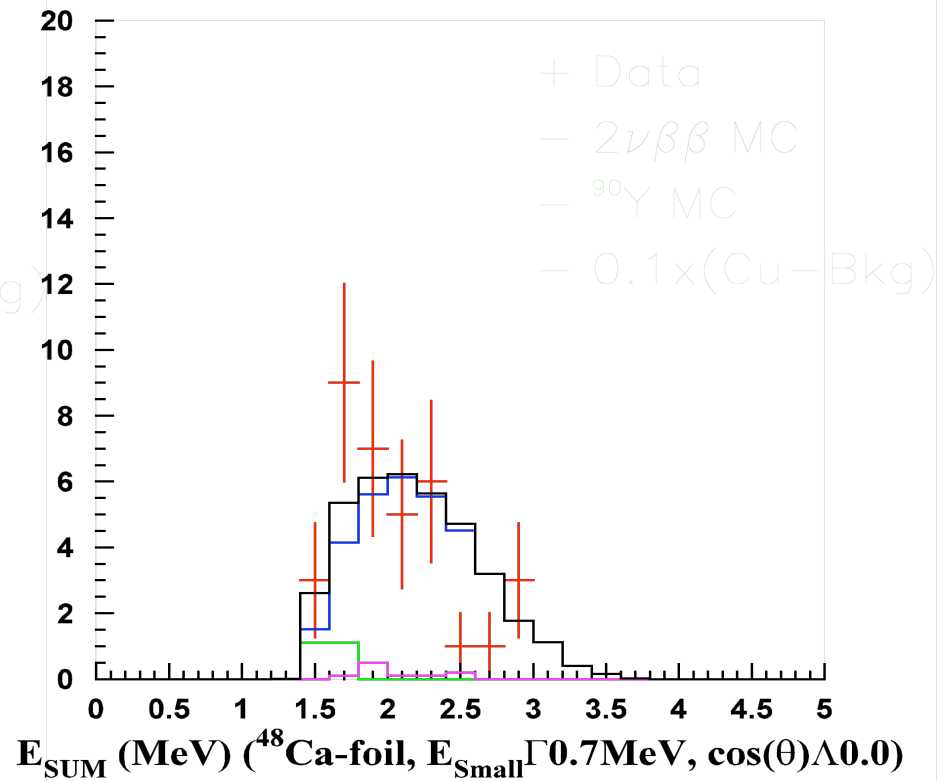
+ Data
 - $2\nu\beta\beta$ MC
 - ^{90}Y MC
 - $0.1 \times (\text{Cu} - \text{Bkg})$

$T_{1/2} = 3.4 \pm 1.2 \cdot 10^{19}$ years

result 35 events background 2



$E_{\text{small}} > 0.2 \text{ MeV}$



$E_{\text{small}} > 0.7 \text{ MeV } \cos(\theta) < 0.0$

Search for $2\beta 0\nu$ decay in NEMO-3

Origin of Background at high energy



Two natural isotopes which have the greatest Q_β values > 3 MeV:

$$^{214}\text{Bi} : Q_\beta \approx 3.27 \text{ MeV}$$

$$^{208}\text{Tl} : Q_\beta \approx 4.99 \text{ MeV}$$

Design NEMO-3 detector for 10 kg:

$$\begin{cases} ^{214}\text{Bi} \text{ in source foils} < 0.3 \text{ mBq/kg} \\ ^{208}\text{Tl} \text{ in source foils} < 0.02 \text{ mBq/kg} \end{cases}$$

Total activity of the detector (30 tons) ≈ 1120 Bq

$$\begin{cases} ^{40}\text{K} = 800 \\ ^{214}\text{Bi} = 300 \\ ^{208}\text{Tl} = 20 \end{cases}$$



In the Modane Underground Laboratory:

$$\text{Fast neutron flux } (> 1 \text{ MeV}): \quad 3.5 \pm 1.5 \quad 10^{-6} \quad \text{n.cm}^{-2}\text{s}^{-1}$$

$$\text{Thermal neutron flux } (\sim 0.025 \text{ eV}): \quad 1.6 \pm 0.1 \quad 10^{-6} \quad \text{n.cm}^{-2}\text{s}^{-1}$$

How NEMO-3 tags the background

- Electron and positron
- Gamma : 50% efficiency at 1 MeV
Energy Threshold = 30 keV
- Time of Flight : Time Resolution \approx 250 ps at 1 MeV
- e^+/e^- separation with a magnetic field of 25 G
3% confusion at 1 MeV
- Delayed tracks ($<700 \mu\text{s}$) to tag delayed α from Bi^{207}
 $^{214}\text{Bi} \rightarrow ^{214}\text{Po} (164 \mu\text{s}) \rightarrow ^{210}\text{Pb}$

ββ0ν Analysis: Background Measurement

NEMO-3 can measure each component of its background !

➤ **External Background ²⁰⁸Tl (PMTs)**

Measured with (e⁻, γ) external events

↪ ~ 10⁻³ ββ0ν-like events year⁻¹ kg⁻¹ with 2.8 < E₁ + E₂ < 3.2 MeV

➤ **External Neutrons and High Energy gamma**

Measured with (e⁻, e⁻)_{int} events with E₁ + E₂ > 4 MeV

↪ ≲ 0.02 ββ0ν-like events year⁻¹ kg⁻¹ with 2.8 < E₁ + E₂ < 3.2 MeV

Only 2 (e⁻, e⁻)_{int} events with E₁ + E₂ > 4 MeV
observed after 260 days of data (without boron)

{ 4253 keV (26 Mar. 2003)
6361 keV (8 Nov. 2003)

In agreement with expected background

sources	A (μBq/kg) from (e ⁻ , Nγ)	A (μBq/kg) HPGe meas.
¹⁰⁰ Mo	92 ± 18	< 110
metal.		
¹⁰⁰ Mo comp.	115 ± 13	< 100
⁸² Se	316 ± 46	400 ± 100

In agreement with HPGe measurements

➤ **¹⁰⁰Mo ββ2ν decay** T_{1/2} = 7.7 · 10¹⁸ y (SSD)

~ 0.3 ββ0ν-like events year⁻¹ kg⁻¹ with 2.8 < E₁ + E₂ < 3.2 MeV

$\beta\beta 0\nu$ Analysis: Background Measurement

Radon in the NEMO-3 gas of the wire chamber

Due to a tiny diffusion of the radon of the laboratory inside the detector
 $A(\text{Radon})$ in the lab $\sim 15 \text{ Bq/m}^3$

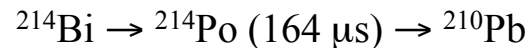
Two independant measurements of radon in NEMO-3 gas

➤ Radon detector at the input/output of the NEMO-3 gas

~ 20 counts/day for 20 mBq/m^3

➤ $(1e^- + 1\alpha)$ channel in the NEMO-3 data:

Delayed tracks ($< 700 \mu\text{s}$) to tag delayed α from ^{214}Po

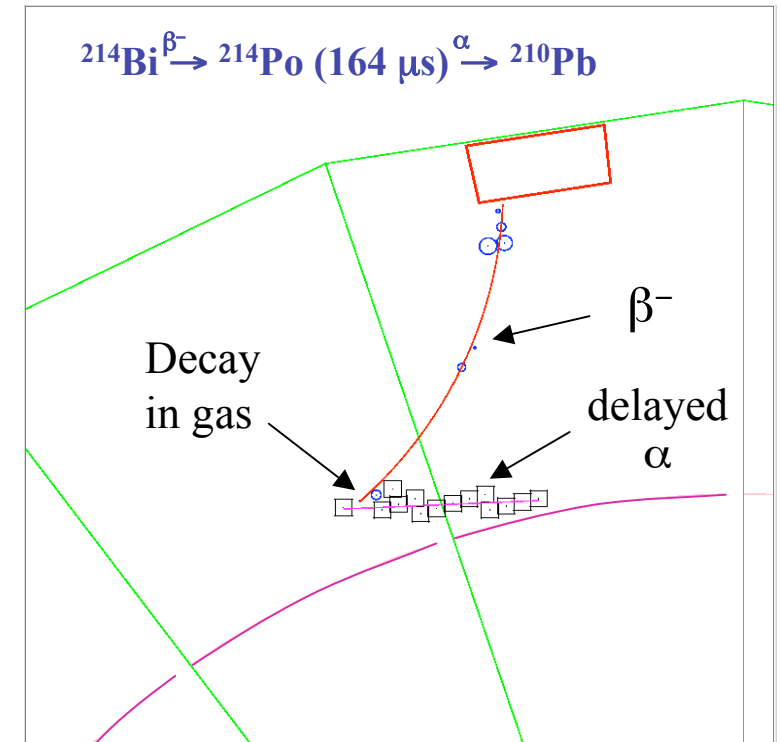


~ 200 counts/hour for 20 mBq/m^3

➡ Good agreement between the two measurements

$A(\text{Radon})$ in NEMO-3 $\approx 20\text{-}30 \text{ mBq/m}^3$

➡ $\sim 1 \beta\beta 0\nu$ -like events/year/kg with $2.8 < E_1 + E_2 < 3.2 \text{ MeV}$



**Radon is the dominant background today
for $\beta\beta 0\nu$ search in NEMO-3 !!!**

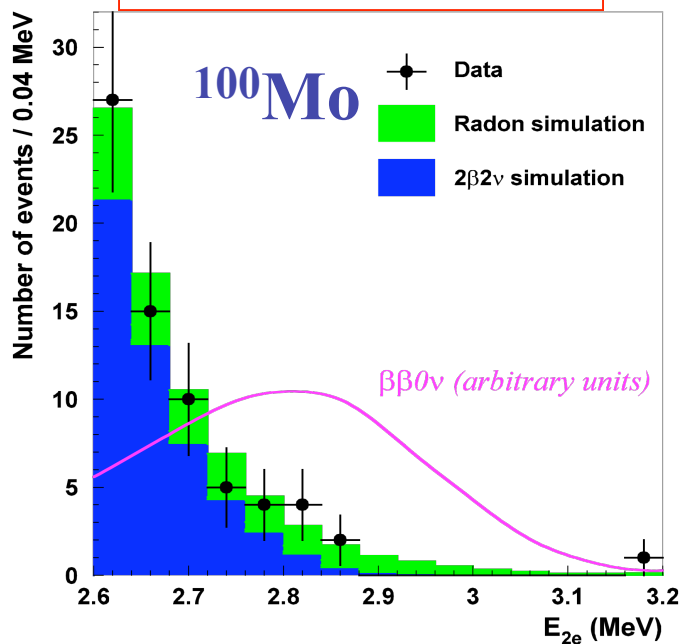
Limit on the effective mass of the Majorana neutrino

Phase 1 (Feb. 2003 – Sept. 2004: 1.08 y of data) with radon bkg
(limits @ 90% CL)

^{100}Mo (6.914 kg)

$T_{1/2}(\beta\beta 0\nu) > 4.6 \cdot 10^{23}$ y

$\langle m_{\nu} \rangle < 0.66 - 2.81$ eV



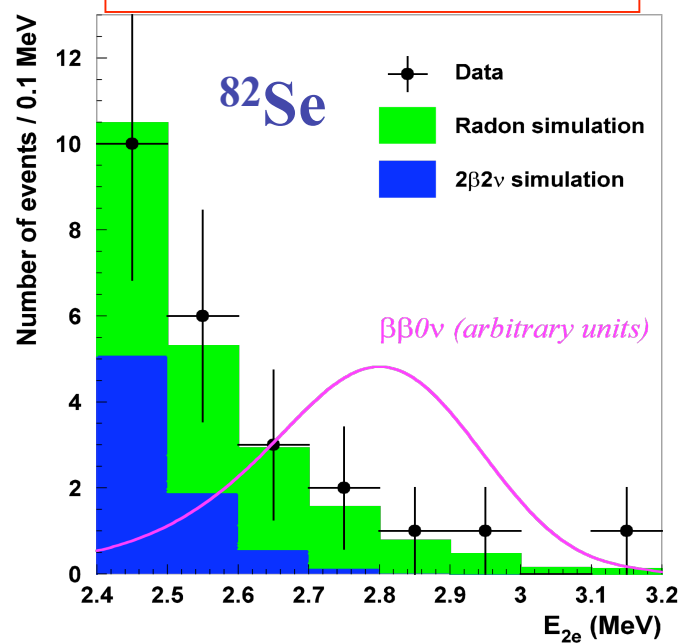
[2.8-3.2] MeV: $\epsilon(\beta\beta 0\nu) = 8\%$
Expected bkg = 8.1 ± 1.3
 $N_{\text{observed}} = 7$ events

Previous limits: $T_{1/2}(\beta\beta 0\nu) > 5.5 \cdot 10^{22}$ y
Ejiri et al. (2001)

^{82}Se (0.932 kg)

$T_{1/2}(\beta\beta 0\nu) > 1.0 \cdot 10^{23}$ y

$\langle m_{\nu} \rangle < 1.75 - 4.86$ eV

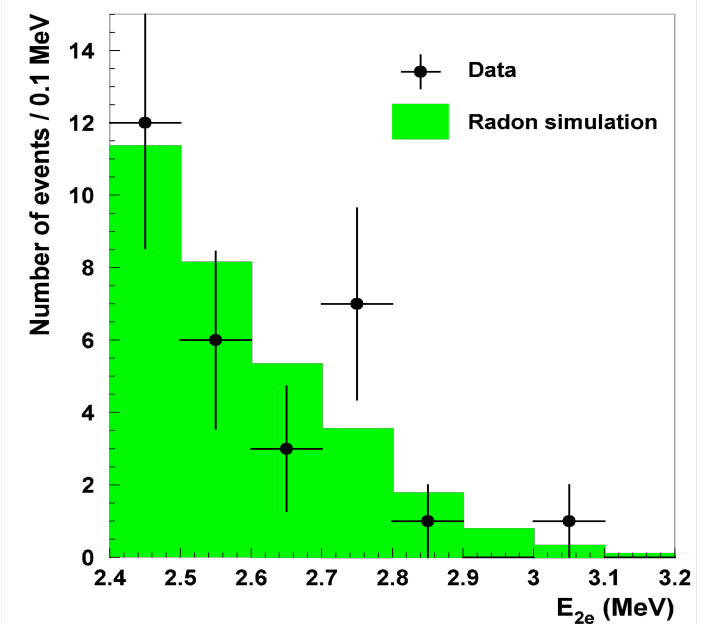


[2.7-3.2] MeV: $\epsilon(\beta\beta 0\nu) = 13\%$
Expected bkg = 3.1 ± 0.6
 $N_{\text{observed}} = 5$ events

Previous limits: $T_{1/2}(\beta\beta 0\nu) > 9.5 \cdot 10^{21}$ y
Arnold et al. (1992)

$\text{Cu} + \text{natTe} + ^{130}\text{Te}$

In agreement with only
Radon bkg expected



Limit on Majoron and V+A

Phase 1 (Feb. 2003 – Sept. 2004: 1.08 y of data) with radon bkg
(limits @ 90% CL)

Limit on Majoron

$$^{100}\text{Mo}: T_{1/2}(\beta\beta 0\nu\text{M}) > 1.8 \cdot 10^{22} \text{ y}$$

$$g_{\text{M}} < (5.3 - 8.5) \cdot 10^{-5} \text{ (best limit)}$$

Simkovic (1999), Stoica (1999)

$$^{82}\text{Se}: T_{1/2}(\beta\beta 0\nu\text{M}) > 1.5 \cdot 10^{22} \text{ y}$$

$$g_{\text{M}} < (0.7 - 1.6) \cdot 10^{-4}$$

Simkovic (1999), Stoica (2001)

Limit on V+A

$$^{100}\text{Mo}: T_{1/2}(\beta\beta 0\nu \text{V+A}) > 2.3 \cdot 10^{23} \text{ y}$$

$$\lambda < (1.5 - 2.0) \cdot 10^{-6}$$

Tomoda (1991), Suhonen (1994)

$$^{82}\text{Se}: T_{1/2}(\beta\beta 0\nu \text{V+A}) > 1.0 \cdot 10^{23} \text{ y}$$

$$\lambda < 3.2 \cdot 10^{-6}$$

Tomoda (1991)

Radon was the dominant background for $\beta\beta 0\nu$ search in NEMO-3

Radon in the NEMO-3 gas of the wire chamber

Due to a tiny diffusion of the radon of the laboratory inside the detector
 $A(\text{Radon})$ in the lab $\sim 15 \text{ Bq/m}^3$

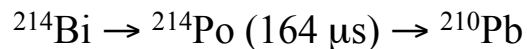
Two independent measurements of radon in NEMO-3 gas

➤ Radon detector at the input/output of the NEMO-3 gas

~ 20 counts/day for 20 mBq/m^3

➤ $(1e^- + 1\alpha)$ channel in the NEMO-3 data:

Delayed tracks ($< 700 \mu\text{s}$) to tag delayed α from ^{214}Po

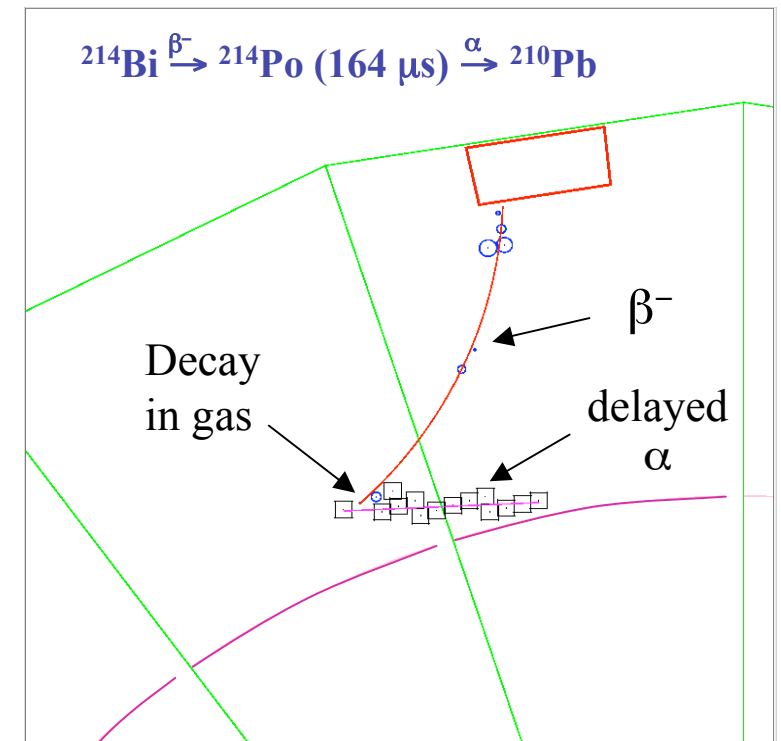


~ 200 counts/hour for 20 mBq/m^3

➡ Good agreement between the two measurements

$A(\text{Radon})$ in NEMO-3 $\approx 20\text{-}30 \text{ mBq/m}^3$

➡ $\sim 1 \beta\beta 0\nu$ -like events/year/kg with $2.8 < E_1 + E_2 < 3.2 \text{ MeV}$



Free-Radon Purification System 1/2

May 2004 : Tent surrounding the detector



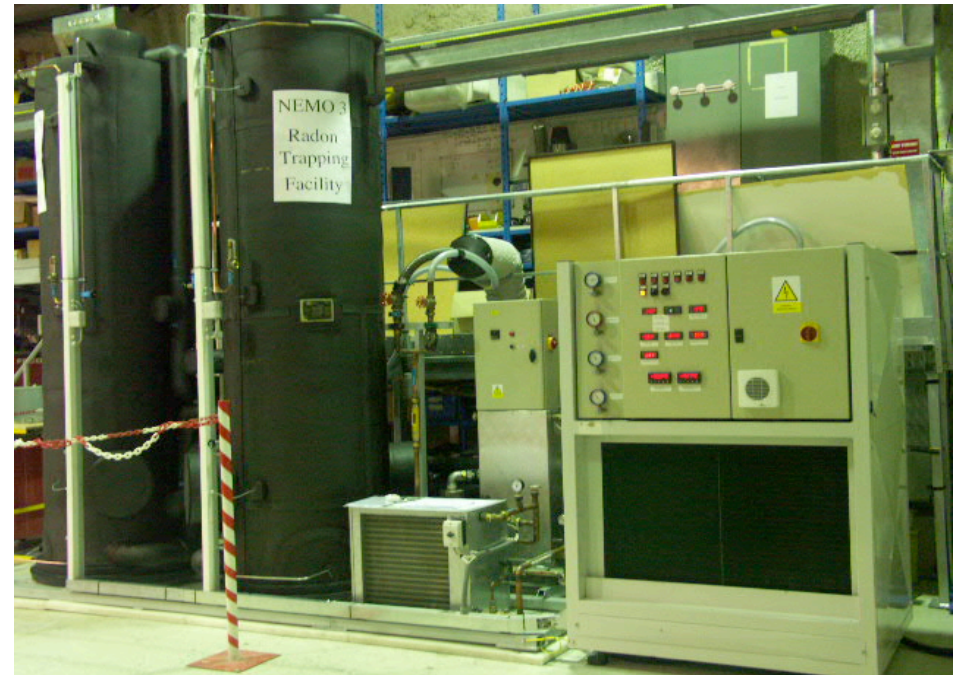
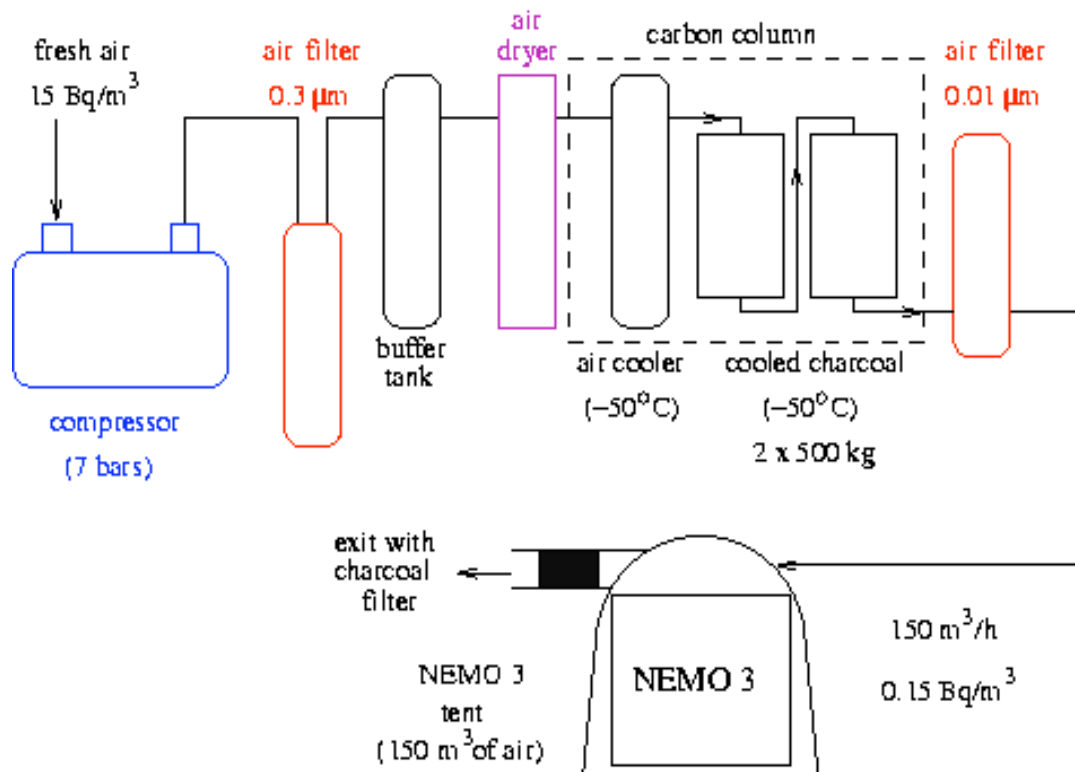
Free-Radon Air factory

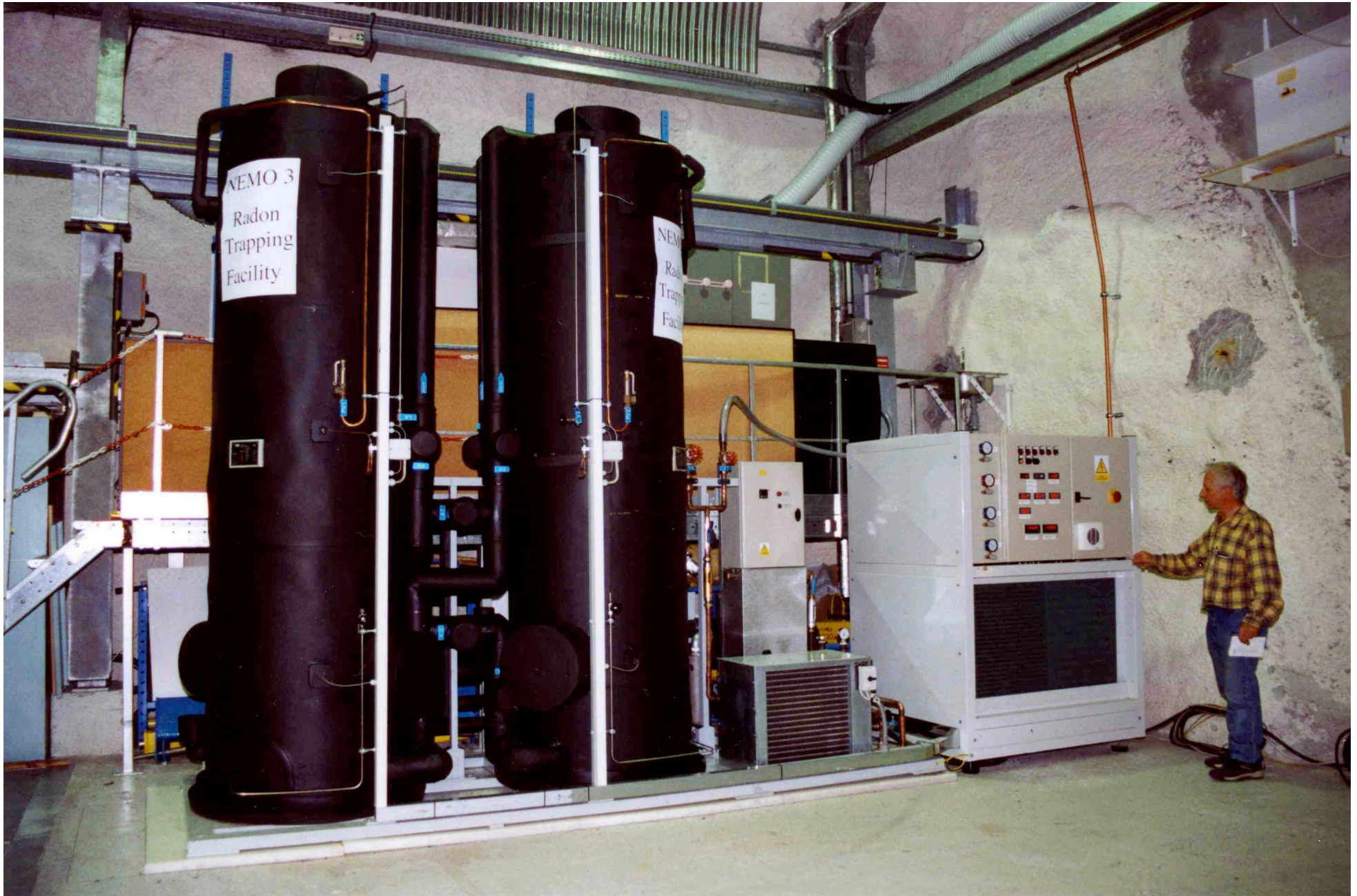
Starts running Oct. 4th 2004
in Modane Underground Lab.

1 ton charcoal @ -50°C, 7 bars

Activity: $A(^{222}\text{Rn}) < 15 \text{ mBq/m}^3$!!!

Flux: $125 \text{ m}^3/\text{h}$ a factor 1000





Radon purification of the air surrounding the detector

- Results -

➤ A(Radon) in the lab..... ~15 Bq/m³

➤ A(Radon) in the tent after flushing free-radon air..... ~ 5 Bq/m³



- source of radon inside the tent
- ground and electronics degasing
- 20-21 oct. 2004: the detector has been roughly isolated from the ground and electronics



➤ A(Radon) in the upper part, surrounding the detector... ~ 0.15 Bq/m³

air flux: 125 m³/h

**Activity of Radon surrounding the detector
has been reduced by a factor ~ 100**

Radon level inside the detector

- Results -

Level of radon measured inside the wire chamber,
by analysing ($1e^- + 1 \alpha$) channel in the NEMO-3 data

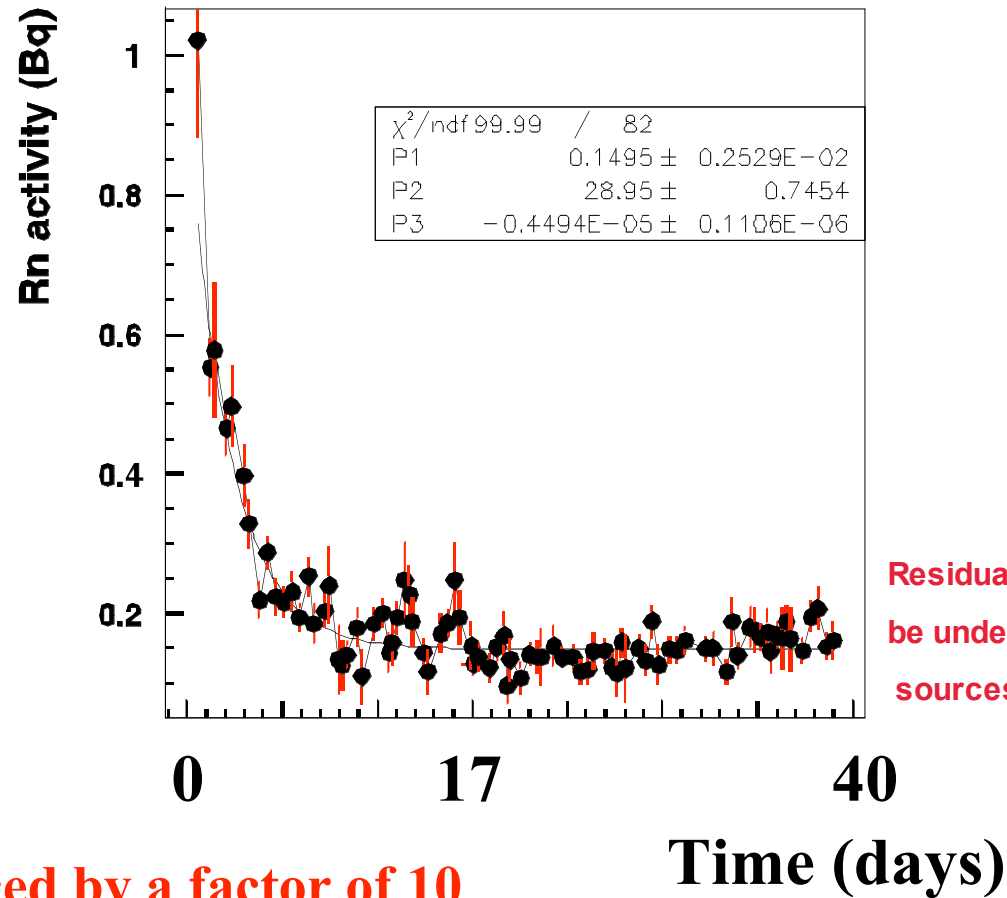
Without tent: $A \sim 1.5$ Bq

After flushing radon-free air
inside the tent: $A \sim 0.15$ Bq



Radon level reduced by a factor of 10

Thanks a lot to S.K especially



M.Nakahata,S.Tasaka

NEMO-3 Expected sensitivity

Background

External Background: negligible

Internal Background: ^{208}Tl : 60 $\mu\text{Bq/kg}$ for ^{100}Mo
300 $\mu\text{Bq/kg}$ for ^{82}Se

^{214}Bi : < 300 $\mu\text{Bq/kg}$

$\sim 0.1 \text{ count kg}^{-1} \text{ y}^{-1}$ with $2.8 < E_1 + E_2 < 3.2 \text{ MeV}$

$\beta\beta 2\nu$ ^{100}Mo : $T_{1/2} = 7.14 \cdot 10^{18} \text{ y}$

$\sim 0.3 \text{ count kg}^{-1} \text{ y}^{-1}$ with $2.8 < E_1 + E_2 < 3.2 \text{ MeV}$



in 2009 after 5 years of data

6914 g of ^{100}Mo $T_{1/2}(\beta\beta 0\nu) > 4 \cdot 10^{24} \text{ y}$ (90% C.L.)
 $\langle m_\nu \rangle < 0.2 - 1.3 \text{ eV}$

932 g of ^{82}Se $T_{1/2}(\beta\beta 0\nu) > 8 \cdot 10^{23} \text{ y}$ (90% C.L.)
 $\langle m_\nu \rangle < 0.6 - 1.7 \text{ eV}$

Prospects

- We live a « prenatal period » for the double beta decay experiments ≈ 10 kg of isotopes NEMO3 ≈ 0.2 eV ^{100}Mo with a small background
CUORECINO ≈ 0.2 eV ^{130}Te , the background would be improved soon
 $^{76}\text{Ge} \approx 0.3$ eV with backgrounds
- NEMO3 phase 2 : a zero background exp. with 10 kg ^{82}Se or a dream 10 kg of ^{150}Nd
- ^{76}Ge experiment : a 10 kg exp. backgrounds would be improved: GERDA
- then experiments with at least 100 kg of isotope to reach
 $\langle m_\nu \rangle \approx \text{a few } 10 \text{ meV}$

Projects near future: 10 years

Experiment	isotope	Mass kg	Type of detector	Lab.	Background (counts FWHM Kg.y)	Sensitivity $T_{1/2}(y)$	Limit	Comment
CUORE	^{130}Te	200	Bolometer	Gran Sasso	0.001	3×10^{26}	0.015-0.090	R&D start
GERDA phase I	^{76}Ge	15	Ge Detector	Gran Sasso	0.01	3×10^{25}	0.3-0.9	2006
GERDA phase II	^{76}Ge	100	Ge Detector	Gran Sasso	0.001	2×10^{26}	0.09-0.29	2009
Super NEMO	^{82}Se ou ^{150}Nd	100	Tracko-Calo	?	0.002	2×10^{26}	0.03-0.06	R&D 2005→ 2007
Majorana	^{76}Ge	500	Ge Detector	?	0.01	4×10^{27}	0.034-0.039	for 10 years of running
EXO	^{136}Xe	200 1000	TPC	WIPP (US)	0.015 0-0.0018	2×10^{26} 8.3×10^{26}	0.39-1.2 0.0510.14	Start 2005 ?
MOON	^{100}Mo	10 1000	Tracko-calo	?		10^{27}	0.2-0.3 0.03	R&D in progress
CANDLES	^{48}Ca	0.180	CaF_2 (200kg)	Otho (Japon)	0.3		0.5	start
COBRA	^{130}Te ^{116}Cd	10	CdZnTe	Boulby U.K	0.2, 0.03	1×10^{24}	0.7	R&D in progress
DCBA	^{150}Nd	20	TPC	?			0.05	R&D in progress
XMASS	Xe	800	Calo.	Kamioka				R&D in Progress 100 kg

What we learnt with NEMO3

- to identify and measure all the sources of background
- to build a very low-background detector
- to prove the reliability of the chosen techniques
- to purify $\beta\beta$ isotopes by removing parents of ^{214}Bi , ^{208}Tl
- to remove background due to Radon (recently)
- to develop ultra low background H P Ge detectors



technique can be extrapolated

R&D program approved recently in France

3 years: 2005, 2006, 2007

From NEMO to SuperNEMO

Factor 100 on the $\beta\beta(0\nu)$ period $T_{1/2}$, reach **few 10^{26} years**

Light Majorana neutrino exchange: $\langle m_\nu \rangle \sim 50$ meV

$$T_{1/2}^{0\nu} \text{ (y)} > \frac{\ln 2 \cdot N}{k_{\text{C.L.}}} \cdot \frac{\epsilon}{A} \cdot \sqrt{\frac{m \cdot t}{N_{\text{BDF}} \cdot R}}$$

Detection efficiency ϵ **Mass of isotope $\beta\beta$ (g)** m
Background ($\text{y}^{-1} \cdot \text{g}^{-1} \cdot \text{keV}^{-1}$) $k_{\text{C.L.}}$ **FWHM (keV)** R
Mass number A **Avogadro Number** N
Measurement time (y) t **Constant** $k_{\text{C.L.}} = 1,6$ à 90% C.L.

Mass
~100 kg

Resolution

(FWHM): ~ 7 % at 3 MeV (will be dominated by source foil)
instead of ~ 11 % at 3 MeV for NEMO 3 (dominated by calorimeter)

Efficiency

improvement by a factor 2

Background

internal contaminations in ^{208}Tl and ^{214}Bi to be improved by a factor of 10

SuperNEMO preliminary design

Plane geometry

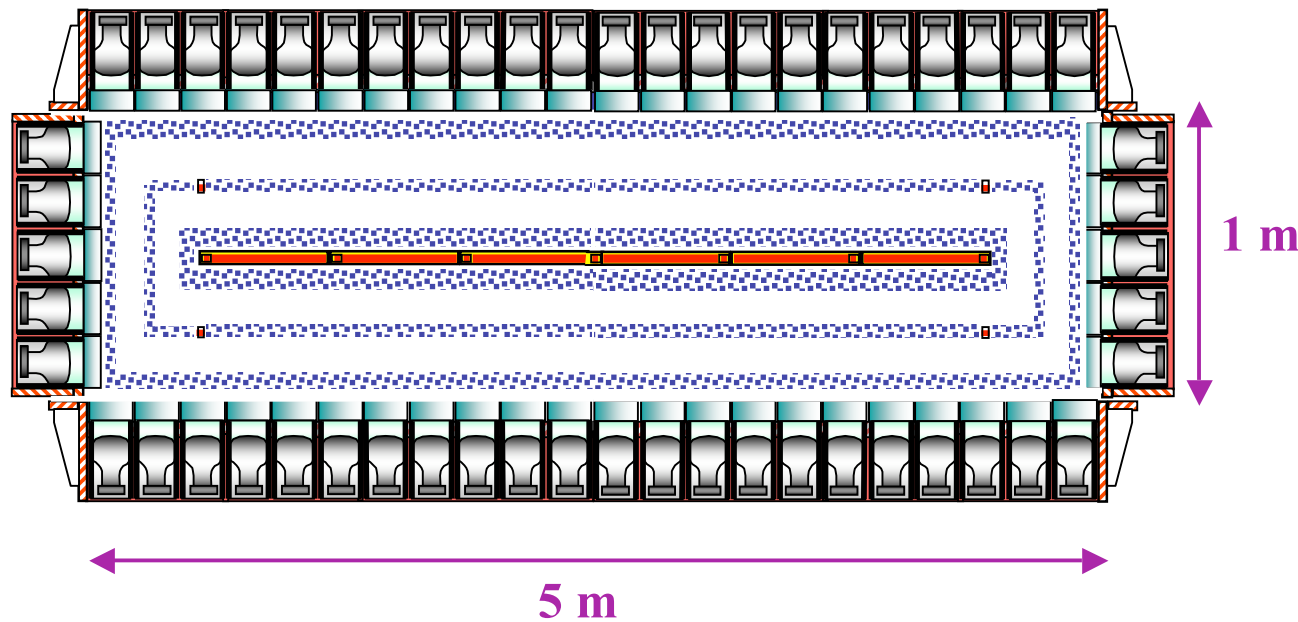
Source (40 mg/cm^2) 12m^2 , tracking volume (~ 3000 channels) and calorimeter (~ 1000 PMT)

Modular ($\sim 5 \text{ kg}$ of enriched isotope/module)

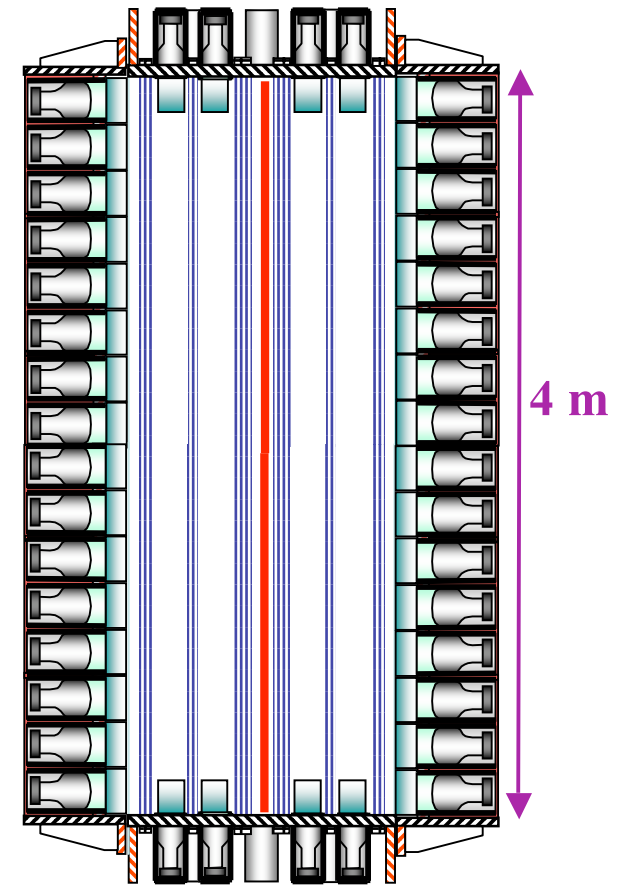
100 kg: 20 modules

$\sim 60\,000$ channels for drift chamber

$\sim 20\,000$ PMT

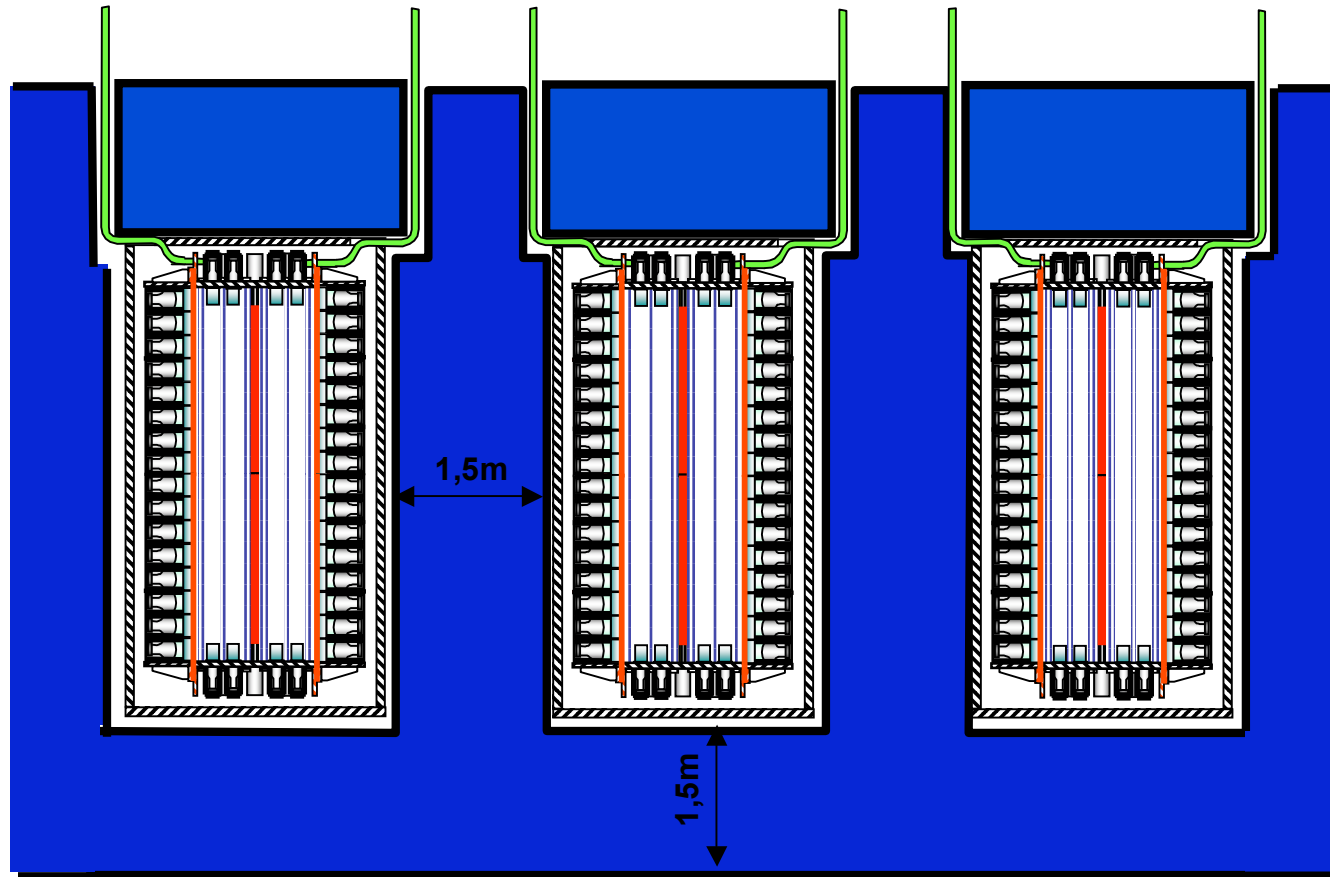


Top view



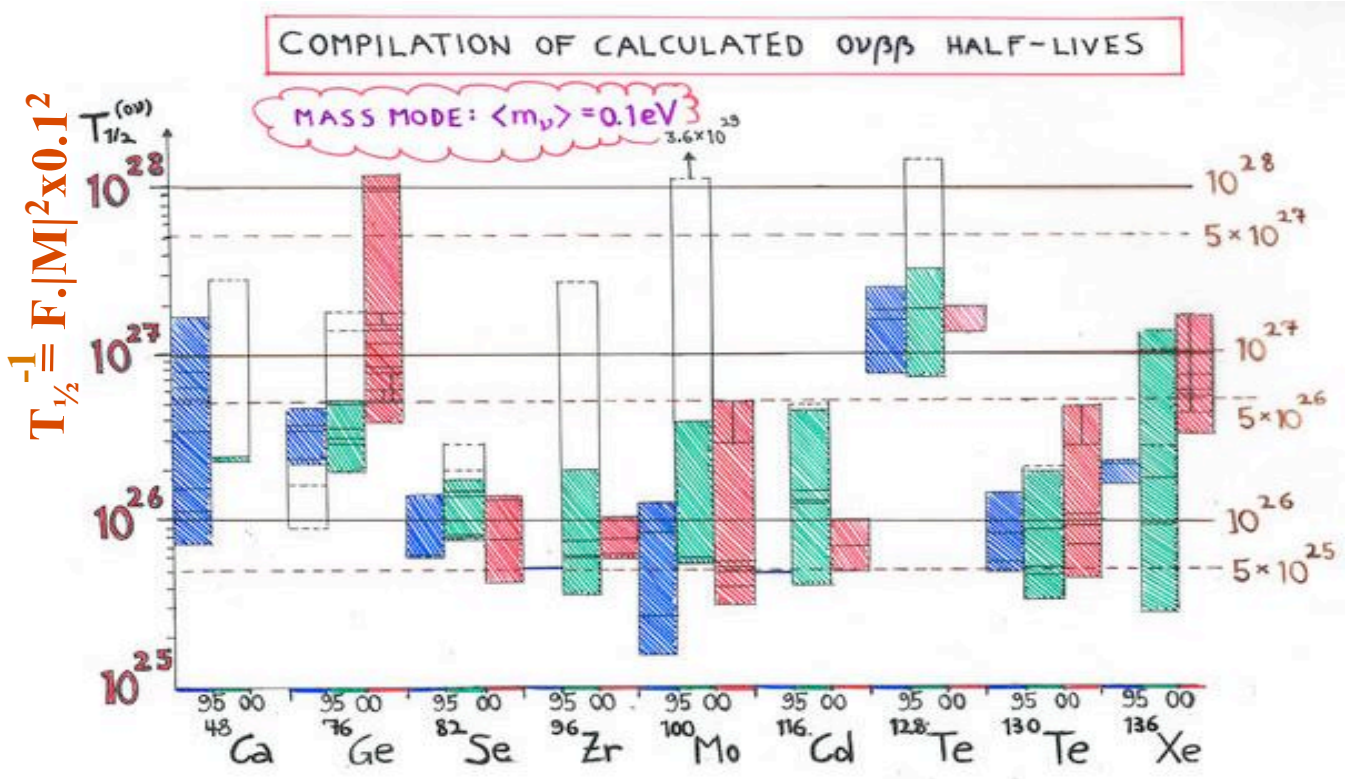
Side view

Water shield



Need of cavity of $\sim 60\text{m} \times 15\text{m} \times 15\text{m}$
Possible in Gran Sasso or in Modane if a new cavity

Choice of the nucleus



Suhonen Neutrino 2004

CONCLUSIONS

- The ranges of values of the $\beta\beta\text{NME}$ is still an open problem
- Need more coherent efforts in computation of $\beta\beta\text{NME}$
- Need new spectroscopic probes



No strong theoretical criteria. Nucleus choice depends on:

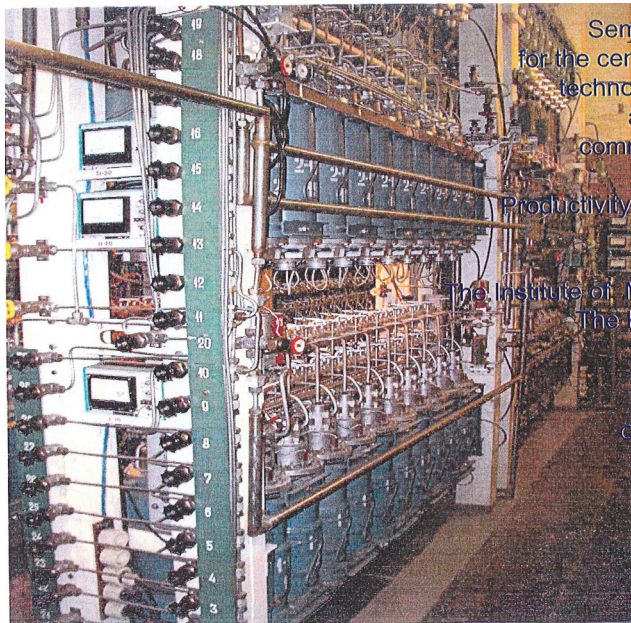
- enrichment possibilities
- experimental technics
- $Q_{\beta\beta}$ value (phase space factor, background)

	$Q_{\beta\beta}$ (MeV)	Isotopic Abundance
$48\text{Ca} \rightarrow 48\text{Ti}$	4.271	0.187
$76\text{Ge} \rightarrow 76\text{Se}$	2.040	7.8
$82\text{Se} \rightarrow 82\text{Kr}$	2.995	9.2
$96\text{Zr} \rightarrow 96\text{Mo}$	3.350	2.8
$100\text{Mo} \rightarrow 100\text{Ru}$	3.034	9.6
$110\text{Pd} \rightarrow 110\text{Cd}$	2.013	11.8
$116\text{Cd} \rightarrow 116\text{Sn}$	2.802	7.5
$124\text{Sn} \rightarrow 124\text{Te}$	2.228	5.64
$130\text{Te} \rightarrow 130\text{Xe}$	2.533	34.5
$136\text{Xe} \rightarrow 136\text{Ba}$	2.479	8.9
$150\text{Nd} \rightarrow 150\text{Sm}$	3.367	5.6

R&D for the source of ^{82}Se

Goal: To be able to produce 100 kg of ^{82}Se with internal contaminations less than $2 \mu\text{Bq/kg}$ in ^{208}Tl and $10 \mu\text{Bq/kg}$ in ^{214}Bi (60 decays/year)

Production: 5kg of ^{82}Se funded by ILIAS
100 kg possible
Development of ICR for enrichment ?



Purification: 2x100 g $^{\text{nat}}\text{Se}$ already processed at INEEL
5kg of ^{82}Se funded by ILIAS
1kg of ^{82}Se 2005
2kg of ^{82}Se 2006
5kg of ^{82}Se 2007

then radiopurity measurements to qualify the process

- 1) H P Ge
- 2) device "à la NEMO"

Thickness: ~250 m² with 40 mg/cm² thickness
(6% (FWHM) at 3 MeV)

Participants: CENBG, LAL, LSCE (France)
ITEP, Kurchatov, JINR (Russia)
INEEL, MHC (USA)

Conclusion

- NEMO3 is running for ≈ 5 years
- R&D program for SuperNEMO 2005, 2006, 2007 a real challenge !
- Coordination started at the European level (ILIAS)
- Neutrinoless double beta decay can be one of the experimental key for understanding neutrino physics: **a long way but promising ?**

Thank you

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